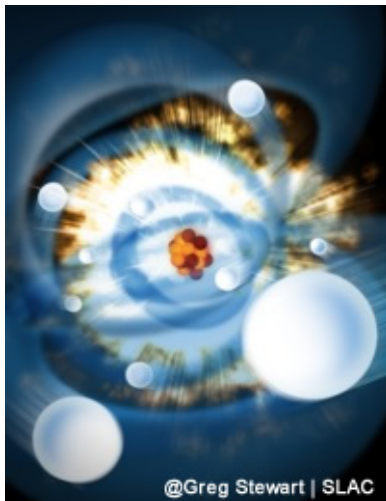


LCLS experiments and x-ray atomic physics

Linda Young



CAMOS Meeting
Keck Center, Washington D.C.
6 April 2011

Outline

- Context & motivation
- Survey of AMO-related LCLS experiments
- Non-resonant high intensity x-ray phenomena
LCLS Experiment 1: Oct 1 - 6, 2009
- Resonant high intensity x-ray processes
LCLS Experiment 5: Oct 29 - Nov 3, 2009
- Outlook/wider impacts/final thoughts

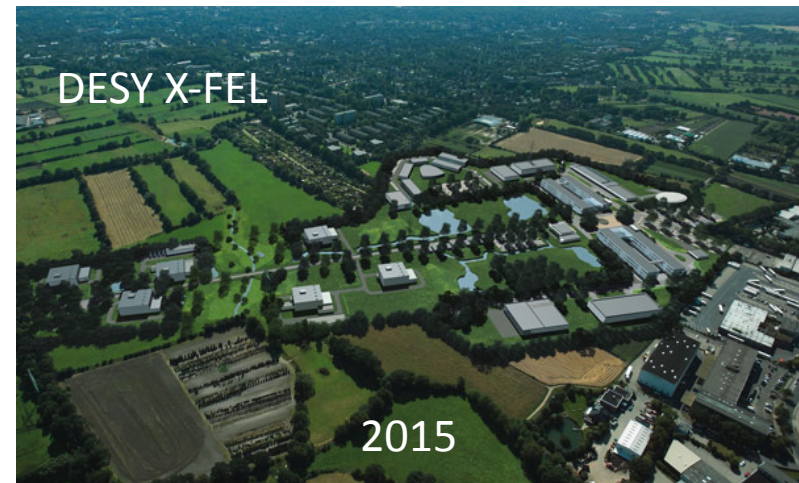
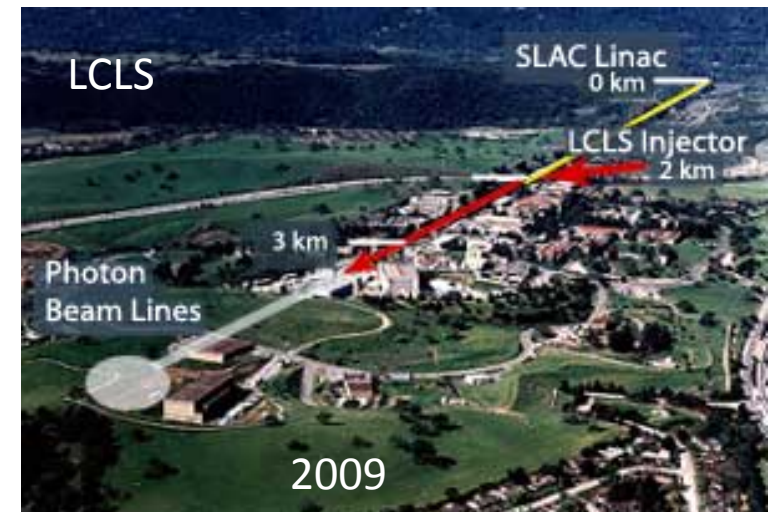


XFELs are here (LCLS) and more are coming soon!

Physics Today, May 2005

X-Ray Free-Electron Laser Projected Parameters			
	LCLS (US)	DESY XFEL (Europe)	SCSS (Japan)
Pulse duration	<230 fs	100 fs	80 fs
Wavelength	1–64 Å	1–15 Å	1–50 Å
Repetition rate	120 Hz	10 Hz	60 Hz
Electron bunches per pulse	1	≤3000	1
Electron beam energy	4–14 GeV	≤20 GeV	≤8 GeV
Photons per pulse ($\times 10^{12}$)	1.2 (at 1.5 Å)	1.2 (at 1 Å)	0.76 (at 1 Å)
Linac length	1 km	2 km	350 m
Estimated cost*	\$379 million	\$1 billion	\$330 million
Estimated start date	2009	2012	2010

*Estimates include varying amounts of instrumentation and different methods of accounting.



Linac Coherent Light Source at SLAC

X-FEL based on last 1-km of existing 3-km linac

Proposed by C. Pellegrini in 1992

1.5-15 Å
(14-4.3 GeV)

Injector (35°)
at 2-km point

Existing 1/3 Linac (1 km)

New e^- Transfer Line (340 m)

X-ray Transport
Line (200 m)

Undulator (130 m)

Near Experiment Hall

Far Experiment
Hall

Lases at 1.5 Å on Apr 10, 2009



UCLA



LLNL



Compare the evolution of high intensity optical and x-ray sources

High-intensity at optical wavelengths

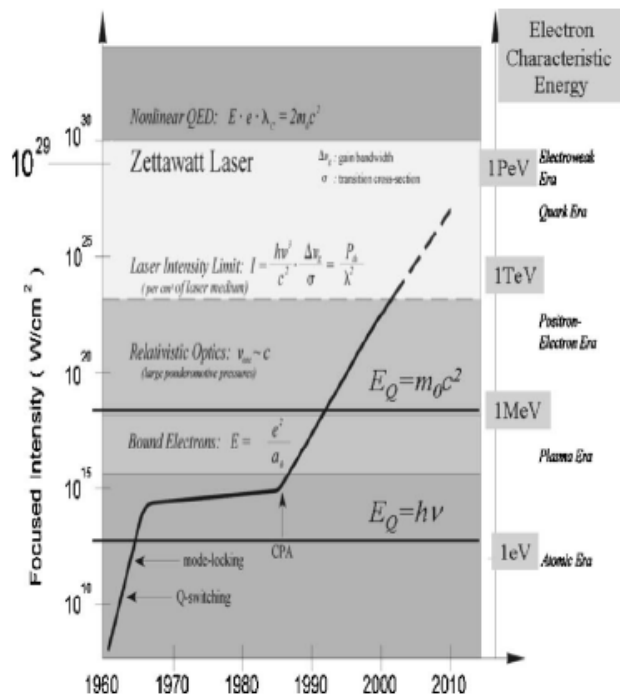
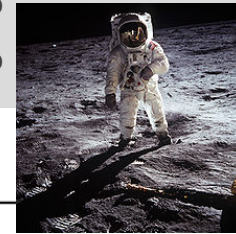
- high harmonic generation
- tabletop coherent x-ray radiation
- attosecond pulses

High-intensity at x-ray wavelengths

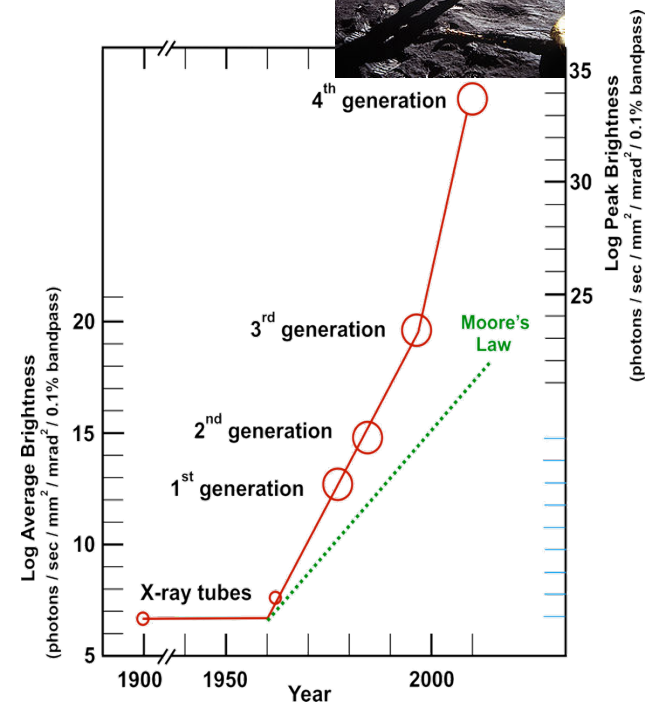
?

?

?



G. Mourou RMP 2006



D. Moncton, George Brown

Contrast optical and x-ray interactions at high intensity

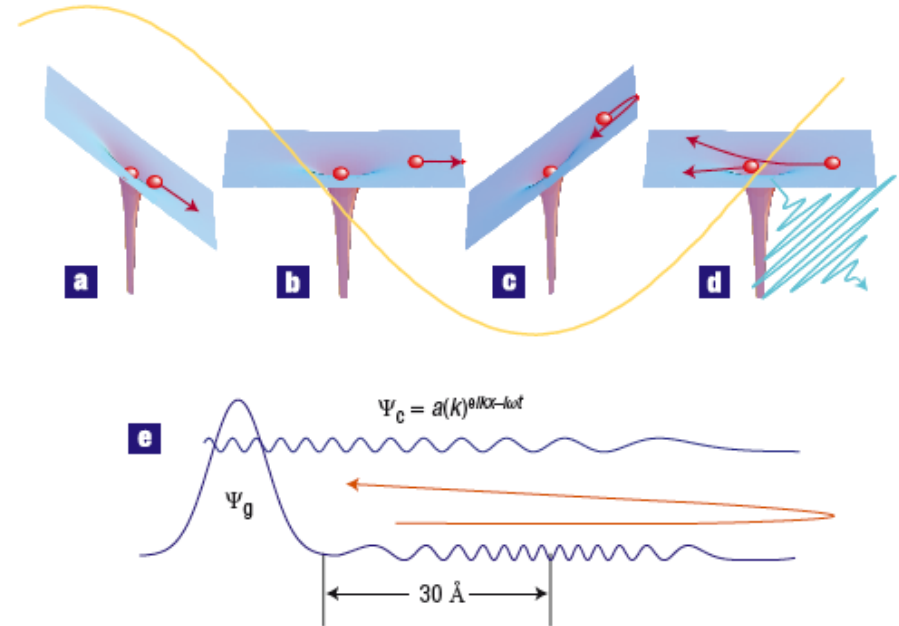
At long wavelengths - laser-driven electron dynamics is dominant
... not so at short wavelengths

electron ponderomotive energy (au)

$$U_p = I/4\omega^2$$

displacement

$$\alpha = E/\omega^2$$



Graphic from Corkum & Krausz
Nature Physics (2007)

Ti:sapphire laser (1.55 eV) PW/cm²
 $U_p \sim 60$ eV & $\alpha \sim 50$ au

LCLS (800 eV) 100 PW/cm²
 $U_p \sim 25$ meV & $\alpha \sim 0.003$ au



Science Drivers for LCLS



AMO: Atomic Molecular and Optical

SXR: Soft X-ray Materials Science

XPP: X-ray Pump-Probe

XCS: X-ray Correlation Spectroscopy

CXI: Coherent X-ray Imaging

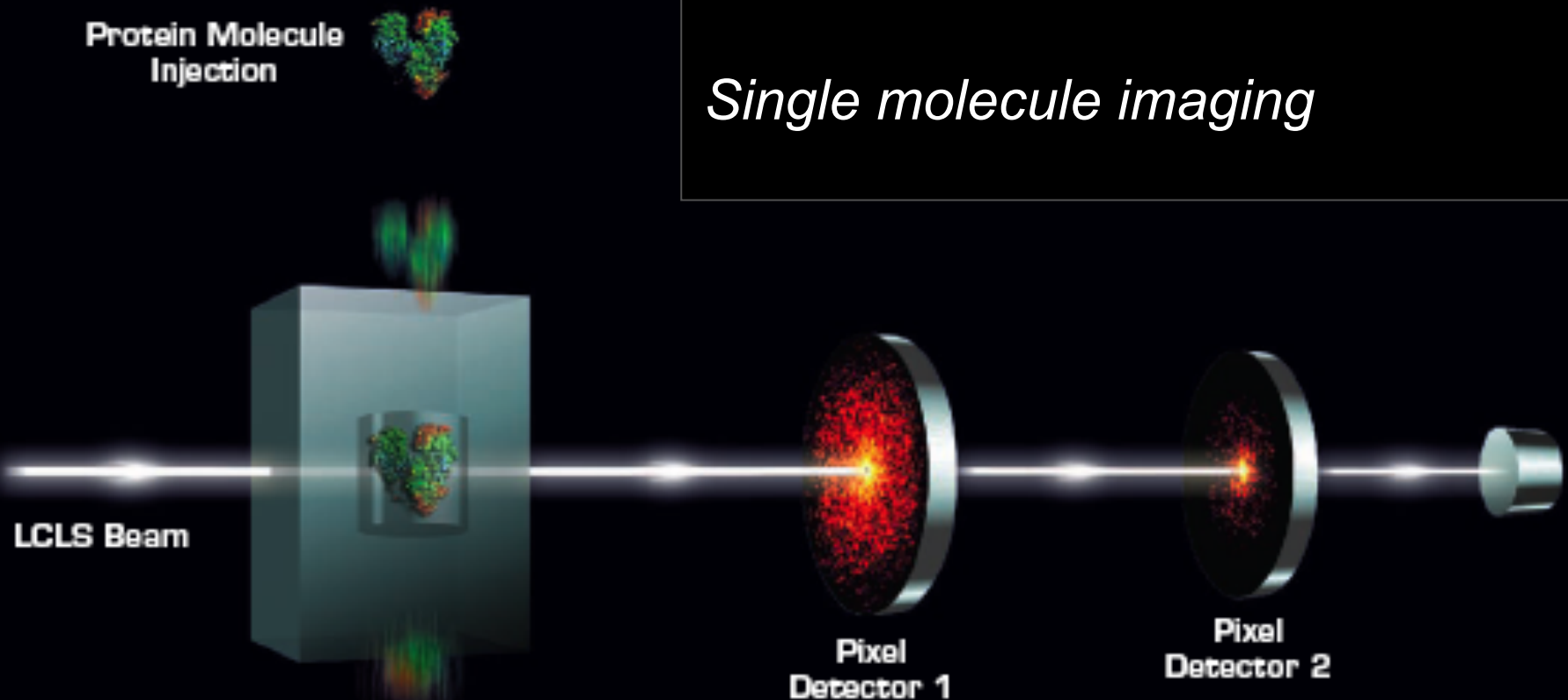
MEC: Materials in Extreme Conditions

AMO

- Understand and control x-ray atom/molecule interactions at ultrahigh x-ray intensity as a foundation for other applications.
- Provide diagnostics of the LCLS radiation



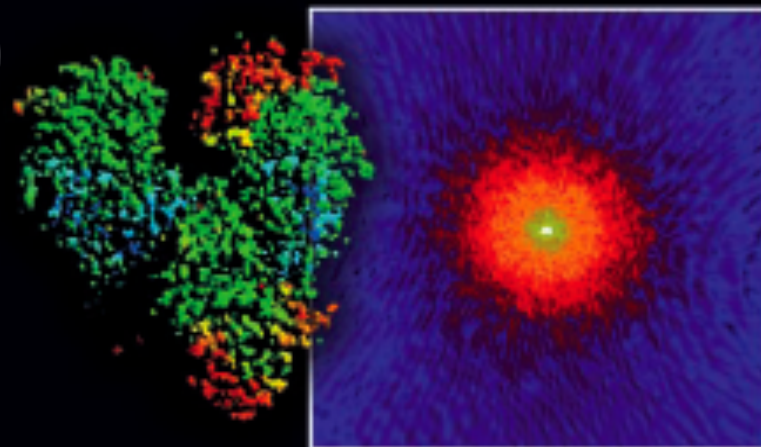
Single molecule imaging



**To Mass
Spectrometer**

**Protein
Molecule**

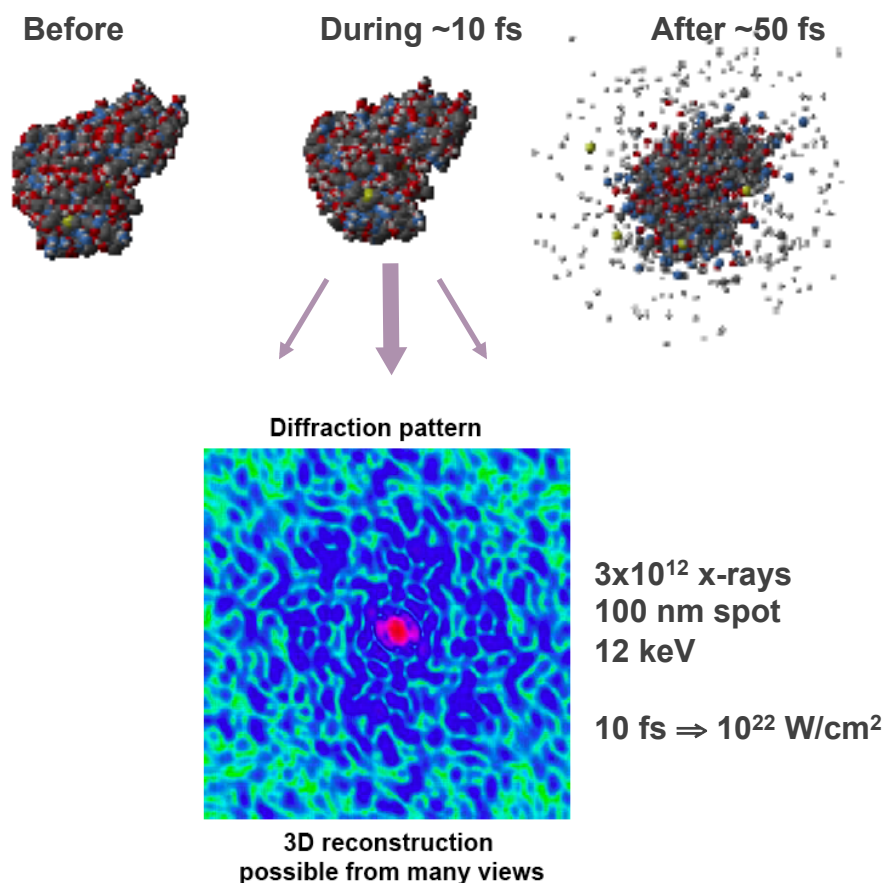
Operating with ultrafast pulses, LCLS will take images of molecules dropped into the x-ray beam. Scientists will merge the series of diffraction patterns of the molecules in many different positions. The resulting three-dimensional reconstruction will reveal the structures of proteins that cannot be crystallized and thus studied any other way.



X-ray Diffraction Pattern

AMO questions at the ultraintense x-ray frontier

- fundamental nature of x-ray damage at high intensity
 - Coulomb explosion
 - electronic damage
 - behavior at 10^{22} W/cm² - 1Å
- nonlinear x-ray processes
 - role of coherence
- quantum control of inner-shell processes



Neutze, Wouts, van der Spoel, Weckert, Hajdu Nature 406, 752 (2000)

AMO-related LCLS publications to date

X-ray Atom Interactions

- *Femtosecond electronic response of atoms to ultraintense x-rays*
L. Young et al., Nature **446**, 56 (2010).
- *Nonlinear atomic response to intense ultrashort x-rays*
G. Doumy et al., Phys. Rev. Lett. **106**, 083002 (2011).

X-ray Molecule Interactions

- *Ultraintense X-Ray Induced Ionization, Dissociation, and Frustrated Absorption in Molecular Nitrogen*
M. Hoener et al., Phys. Rev. Lett. **104**, 253002 (2010).
- *Auger Electron Angular Distribution of Double Core-Hole States in the Molecular Reference Frame*
J. Cryan et al., Phys. Rev. Lett. **105**, 083004 (2010).
- *Double Core-Hole Production in N₂: Beating the Auger Clock*
L. Fang et al., Phys. Rev. Lett. **105**, 083005 (2010).

X-ray / optical laser timing and synchronization

- *Time-resolved pump-probe experiments at the LCLS*
J. M. Glownia et al. Opt. Express **18**, 17621 (2010).



AMO-related LCLS pubs in the works ...

X-ray pulse characterization using two color interactions

- *Femtosecond X-ray Pulse Length Characterization at the LCLS Free Electron Laser*
S. Düsterer M. Meyer et al.
- *Investigation of the time structure of LCLS's few fs x-ray pulses*
R. Kienberger et al.

X-ray Cluster Interactions

- *Explosions of clusters in intense X-ray pulses*
T. Ditmire et al.
- *Ultrafast imaging of X-ray excited clusters*
C. Bostedt, T. Möller et al.

Resonant x-ray interactions: Control of inner-shell processes

- *Modifying Auger Decay with Femtosecond X-ray Pulses*
E. P. Kanter et al.

Electron-Beam Ion Trap: highly-charged ion spectroscopy

Bernitt, Brown, J. Crespo, P. Beiersdorfer, J. Ullrich



LCLS Experiment 1 - Oct 1, 2009

Nature of the electronic response to

10^5 x-rays/ $\text{\AA}^2$
80 - 340 fs
800 - 2000 eV

$\sim 10^{18}$ W/cm 2

Original single molecule imaging parameters, Neutze et al. Nature (2000)

3×10^{12} x-rays/(100 nm) 2 = 3×10^6 x-rays/ $\text{\AA}^2$
10 fs

$\sim 10^{22}$ W/cm 2



Our approach to understanding ultraintense x-ray interactions

- Start with a well-characterized target

Binding energies in neutral neon

2p : ~21 eV

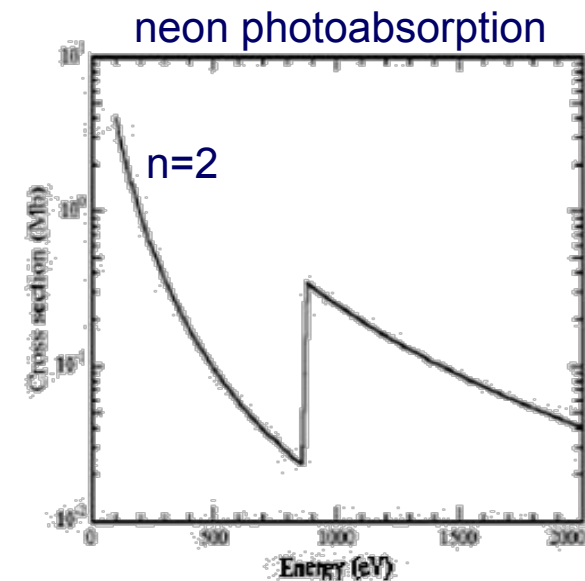
2s : ~48 eV

1s : ~870 eV

Inner-shell excitation

Auger yield 98%

Auger clock - τ_{1s} : 2.4 fs

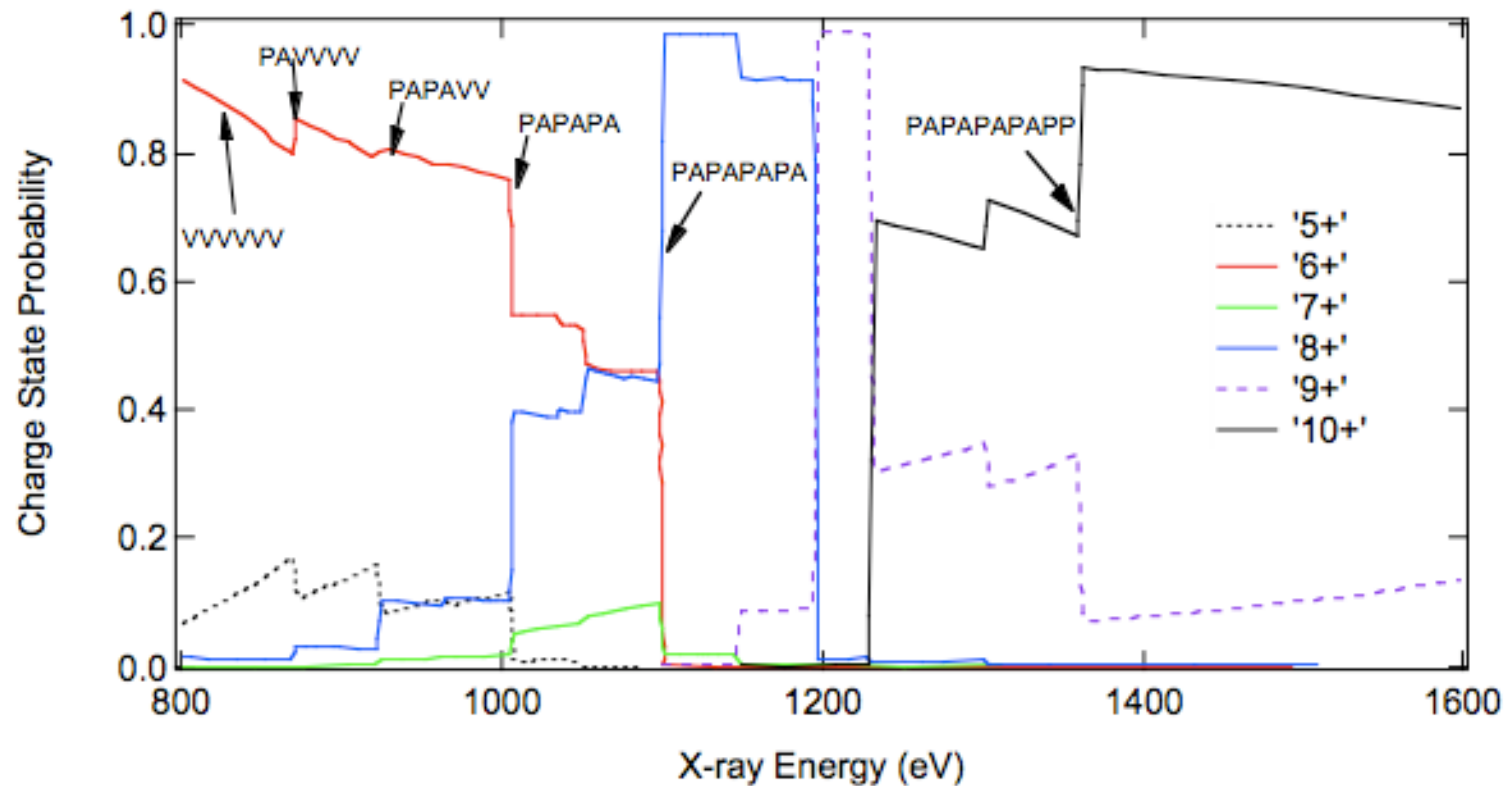


- Probe changes in interaction from outer- to inner-shell between 800-2000 eV



Guided by theory

Theory: Rohringer & Santra, PRA 76, 033416 (2007)

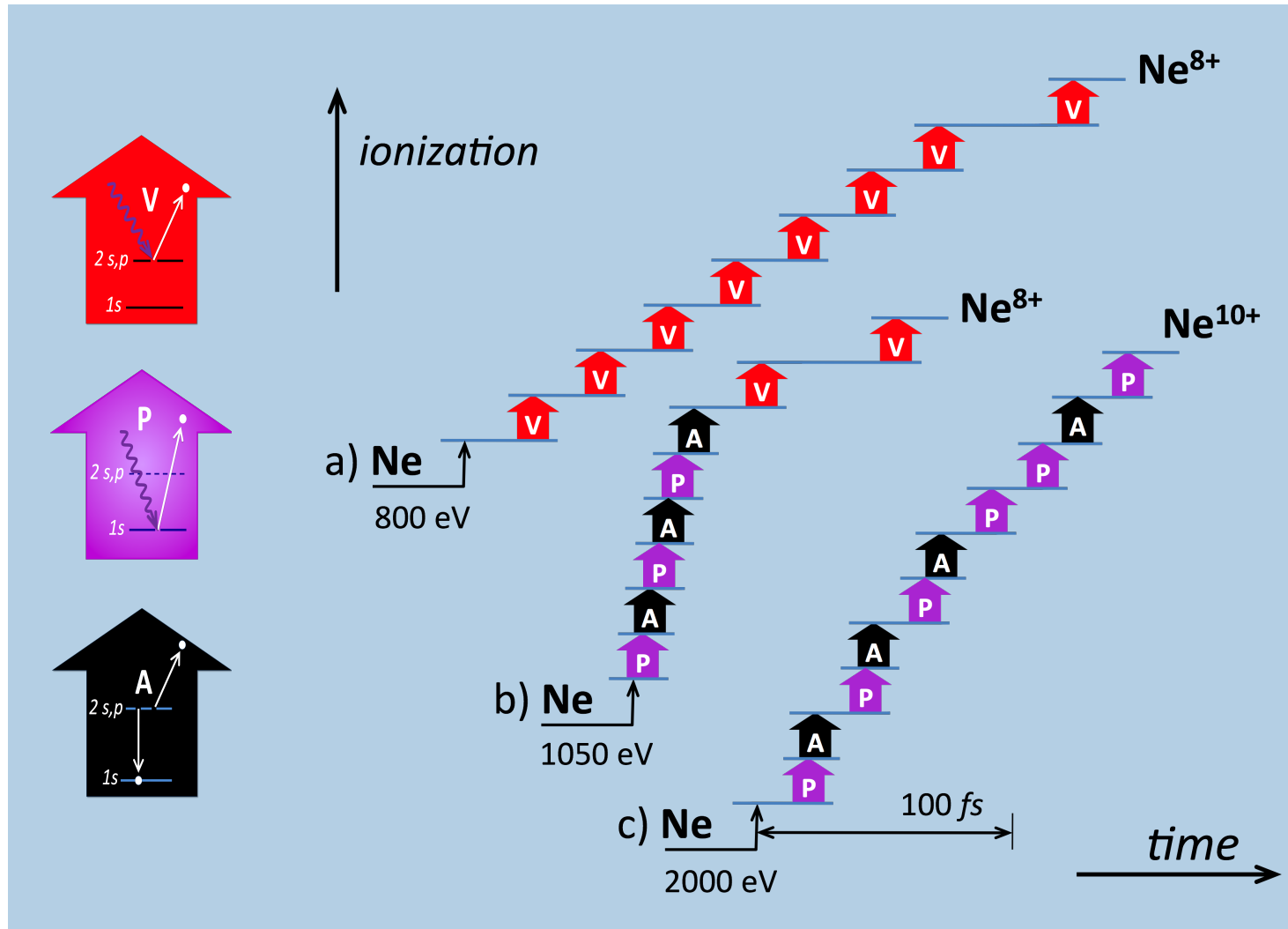


LCLS specs
10¹³ x-rays
230 fs
1 μ m spot

Three target energies: 800 eV, 1050 eV, 2000 eV

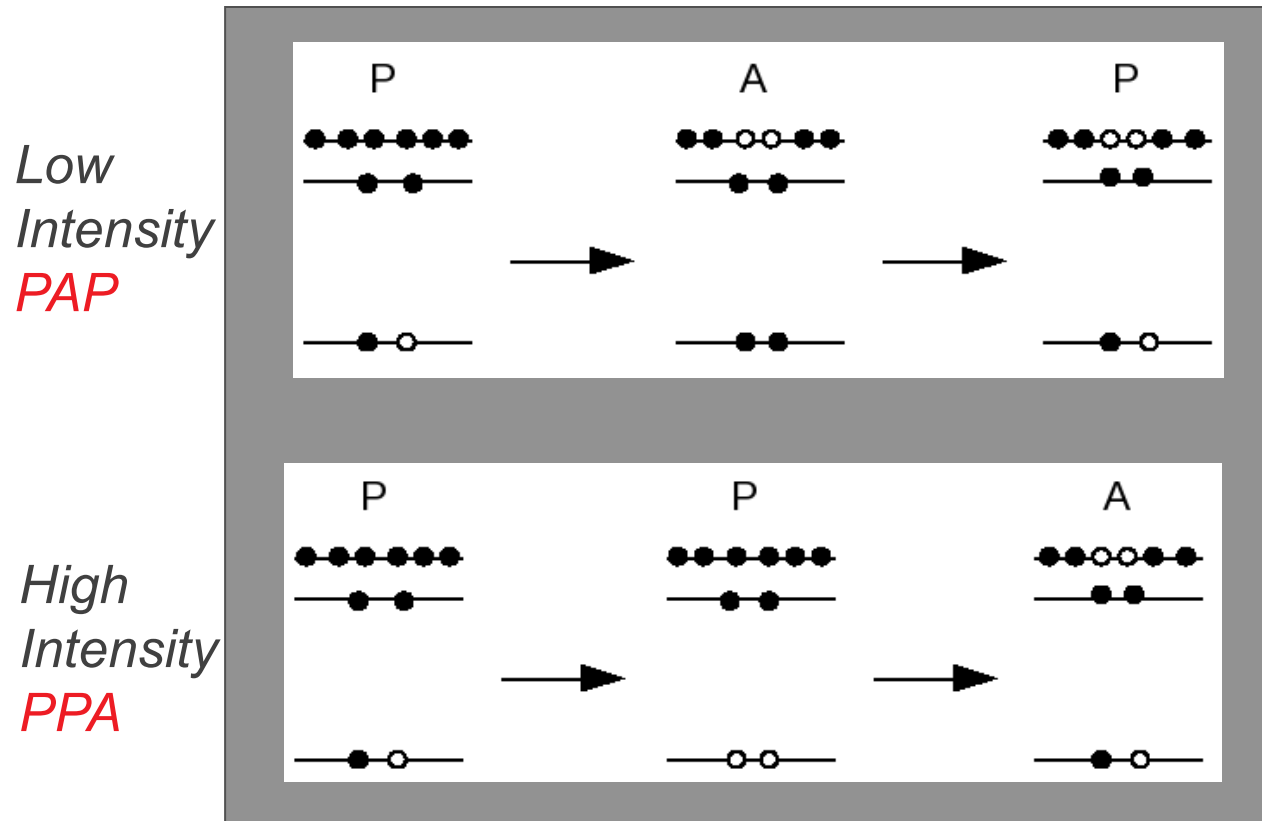


Valence ionization, core ionization and Auger decay



Sequential single photon processes dominate the interaction

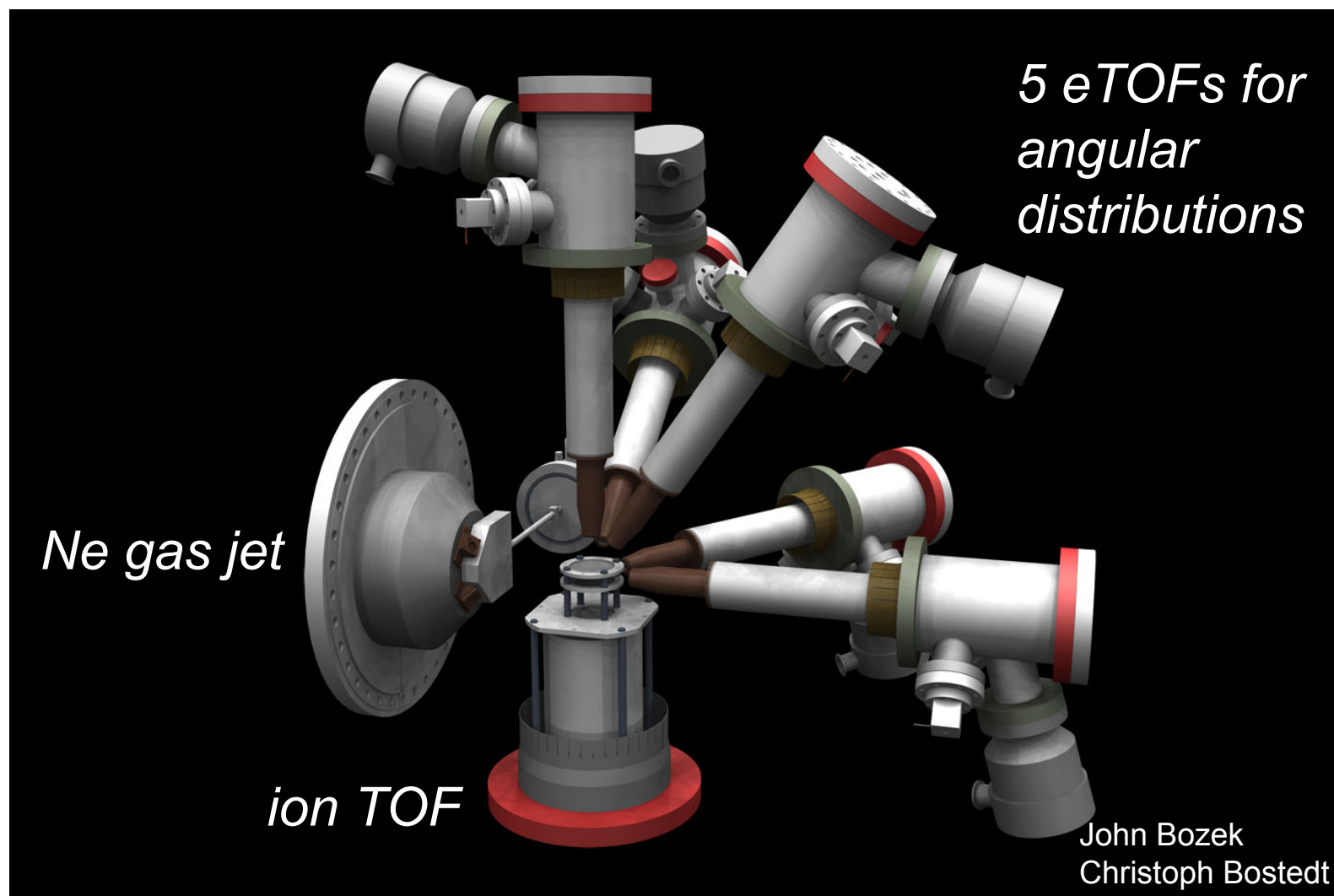
How does one arrive at a particular charge state?



- Hollow atoms produced at high x-ray intensity
- Electron spectroscopy can define the mechanism

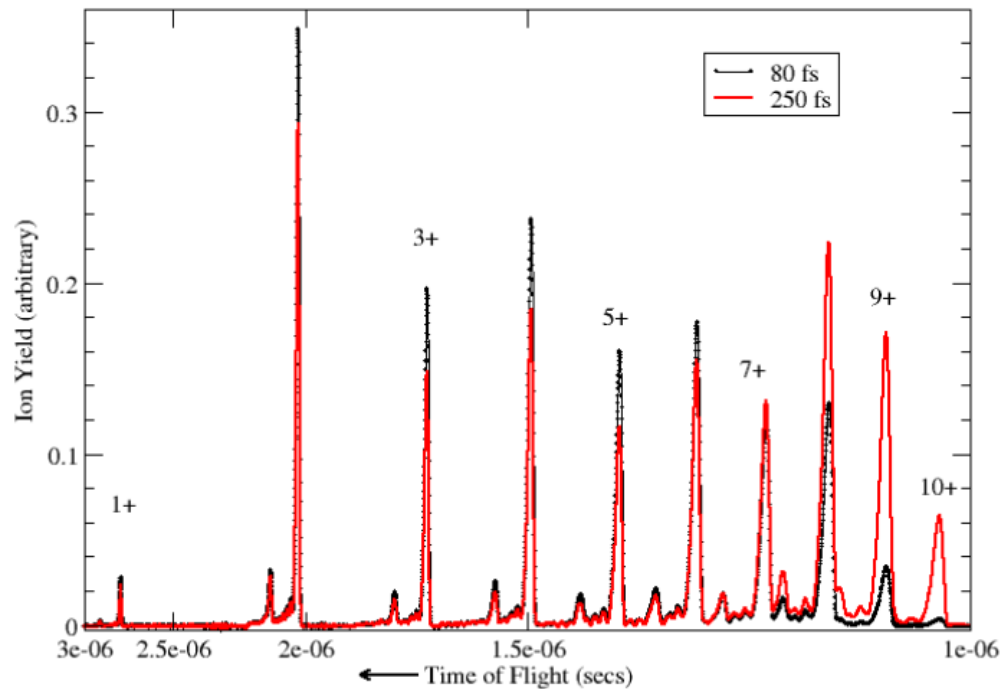


High field physics chamber



Day 1 - two interesting observations

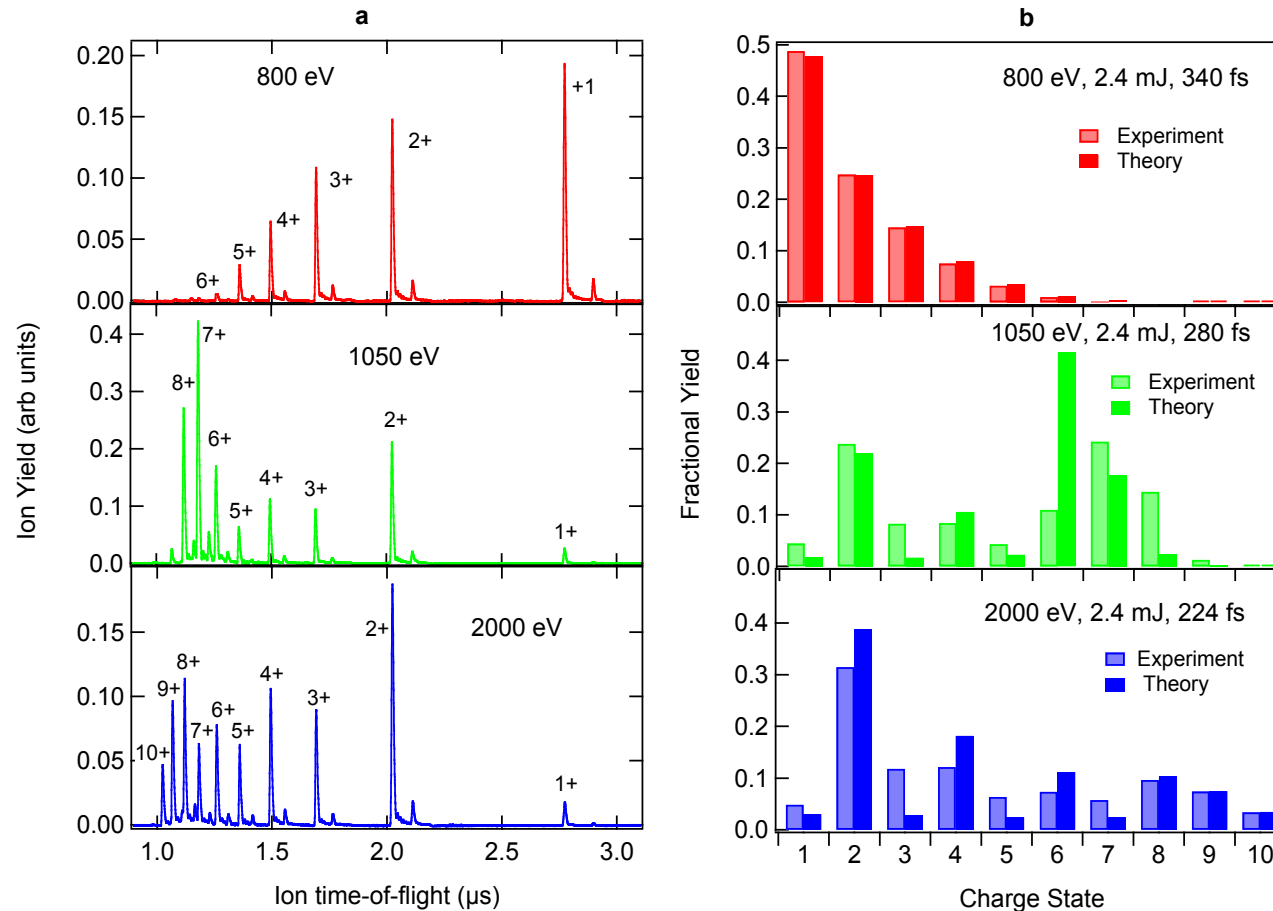
- Single ~ 100 fs pulse at 2000 eV fully strips neon
6-photon, 10-electron process



- Shorter pulses with equal pulse energy & fluence suppress absorption & damage.



Theory can model ultraintense x-ray-induced electronic damage in neon



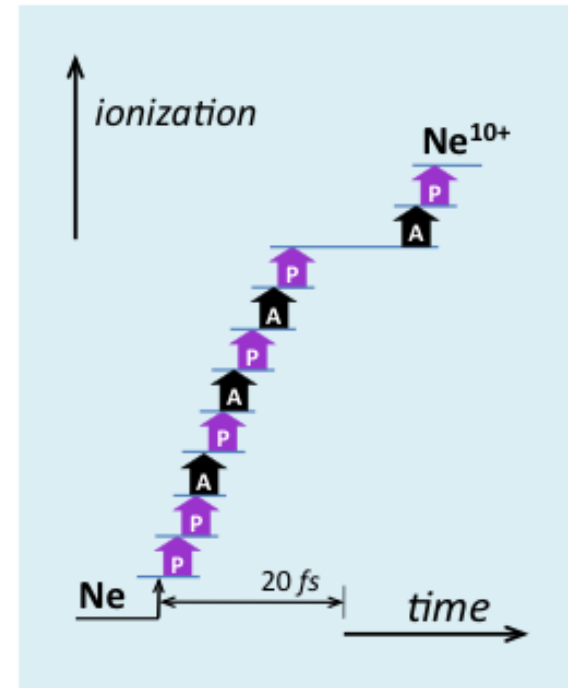
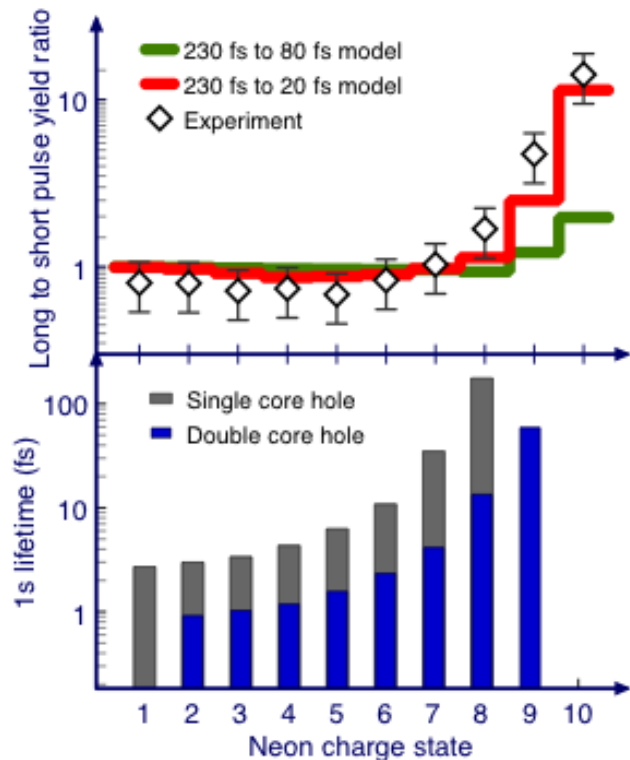
Theory

- Intensity averaged
- Fluence determined by experiment

Consistent with
“measured” pulse
energy and focus.

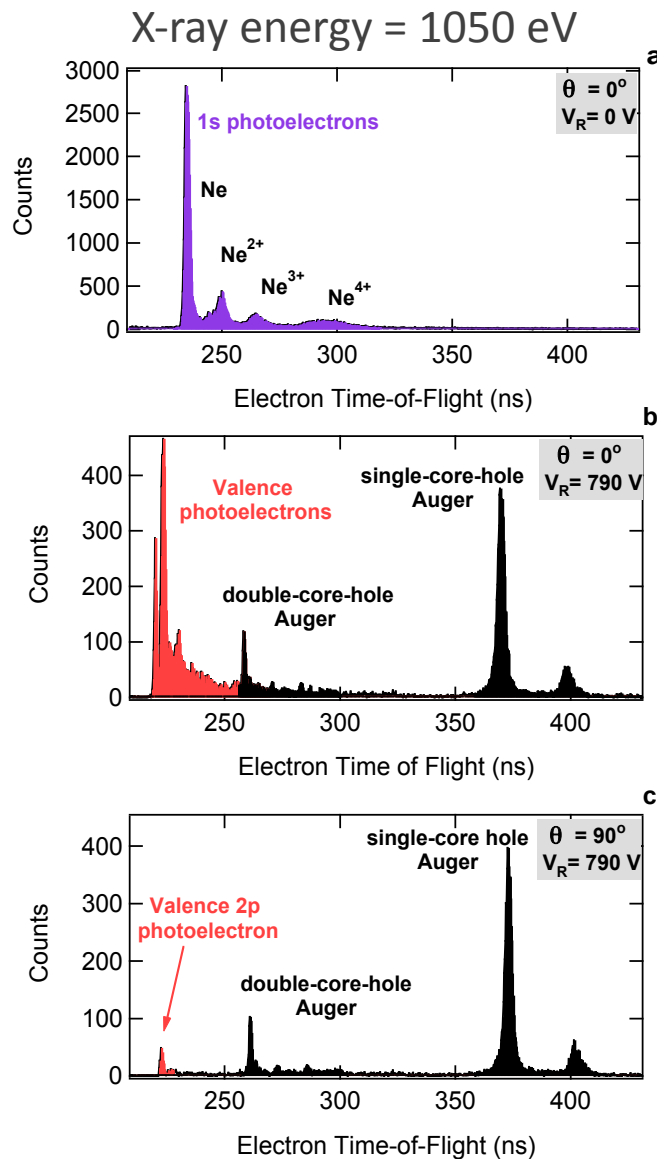
Sang-Kil Son, Robin Santra – refined calcs include shakeoff – G. Doumy et al, PRL 2011)

Atoms become transparent at high x-ray intensity !



- x-ray absorption is due to the presence of 1s electrons
- high x-ray intensities eject 1s electrons rendering the atom transiently transparent
- slowing atomic clocks create transparency at surprisingly long timescales

Electron spectrometers track ionization mechanism

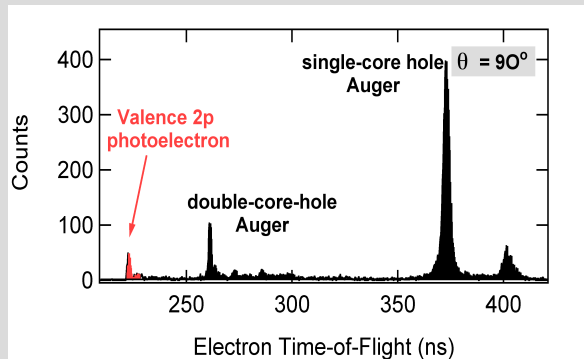


“Slow” 1s photoelectrons along x-ray polarization axis

“Fast” valence photoelectrons and Augers along polarization axis

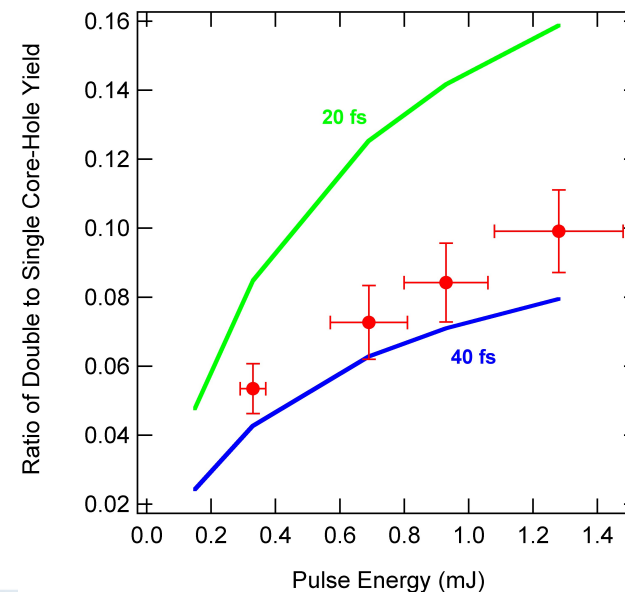
Clean hollow atom signature
double-core-hole Auger
 $\theta = 90^\circ$

Hollow atom production: deliberate, huge and an indicator of x-ray pulse duration

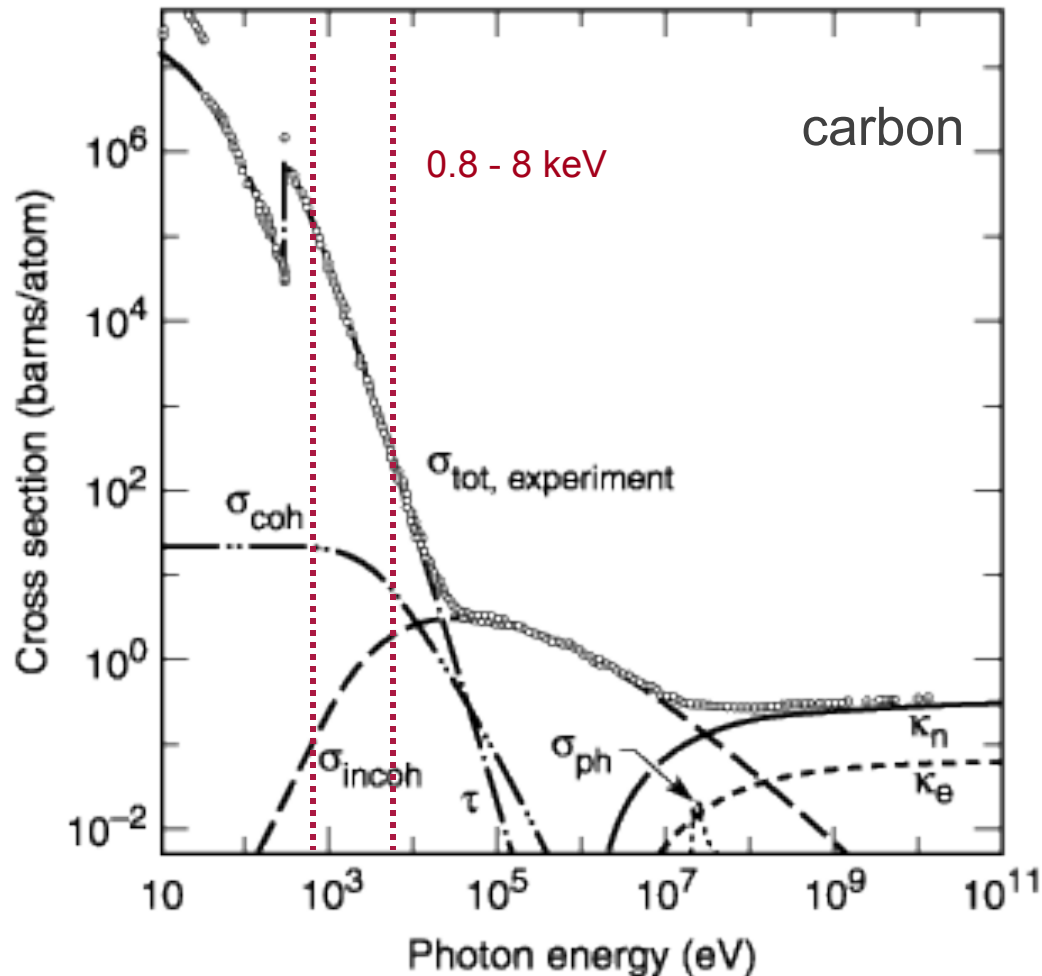


Hollow atom yield
@ LCLS ~10%
@ synchrotron ~0.3%
due to electron correlation

1050 eV,
nominal electron bunch
duration ~80 fs



Absorption vs scattering: normal and hollow atoms



	$\frac{\sigma_{\text{photo}}}{\sigma_{\text{elastic}}}$	$\frac{\sigma_{\text{Compton}}}{\sigma_{\text{elastic}}}$
2 keV	360	0.05
8 keV	20	0.60
8 keV hollow	2	

Impact of hollow atom formation on coherent x-ray scattering

Sang-Kil Son, L Young, R Santra
Phys. Rev A. **83**, 033402 (2011)





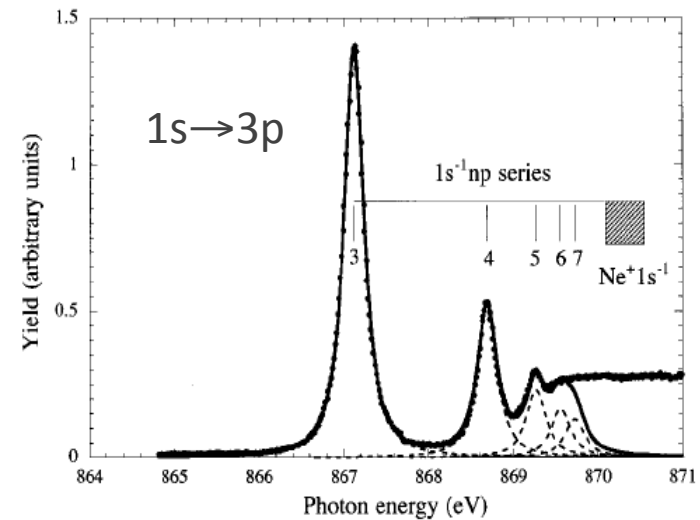
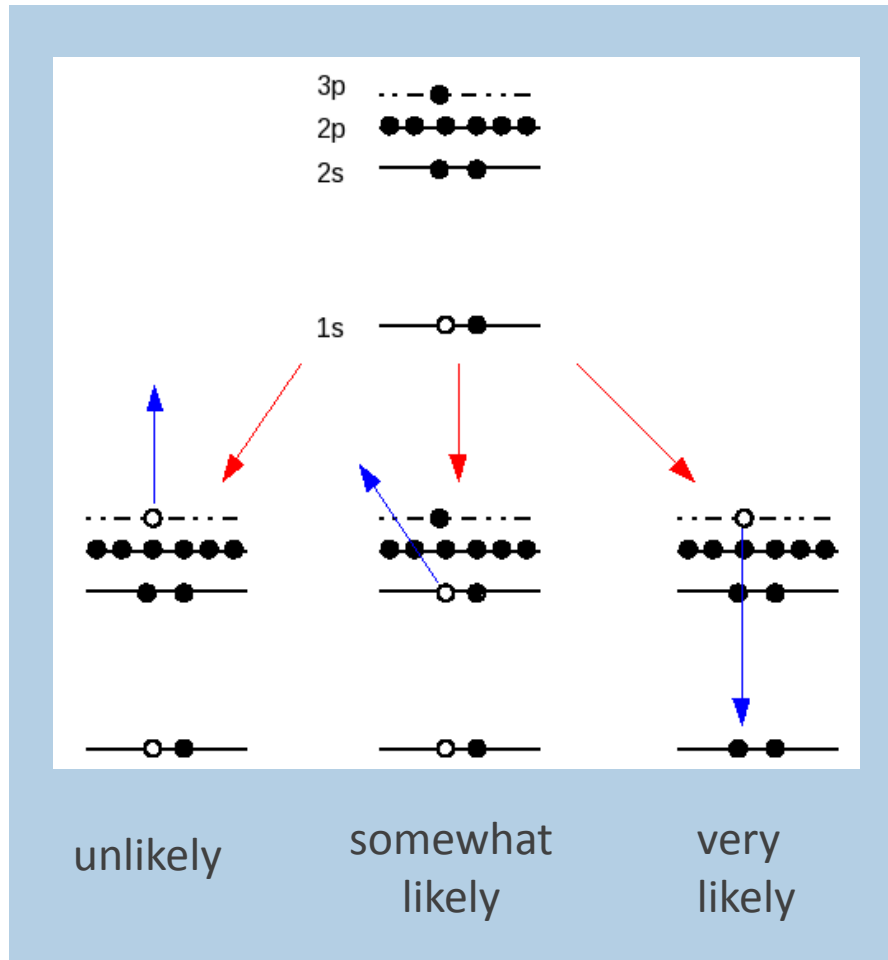
LCLS Experiment 5

Resonant x-ray processes at high intensity



Can we control inner-shell electron dynamics?

“Rabi flopping” may inhibit Auger decay & x-ray damage.



- Strong 1s→3p resonance

$$\mu_{\text{Ne } 1s-3p} = 0.01 \text{ ea}_0$$

$$\tau_{\text{Ne } 1s^{-1}} = 2.4 \text{ fs} = 100 \text{ a.u.}$$

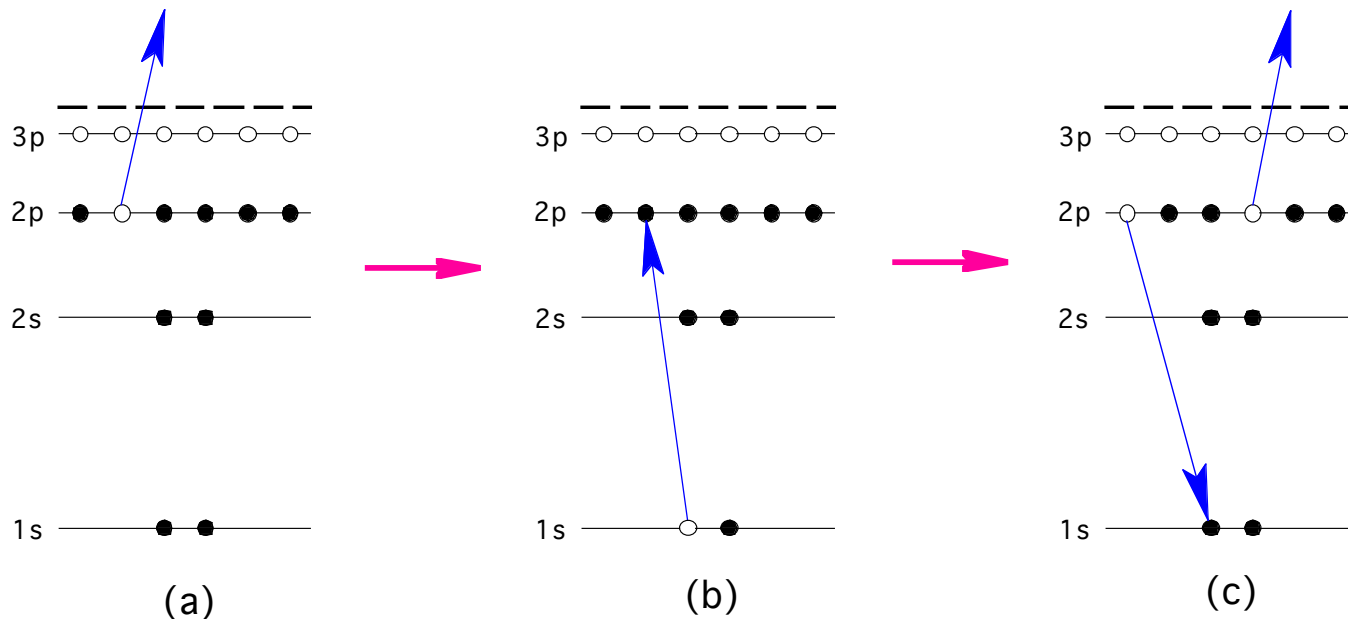
- Rabi flopping possible

$$E_{\text{Ne}} \sim 6.3 \text{ a.u.}$$

$$I_{\text{Ne}} \sim 1.4 \times 10^{18} \text{ W/cm}^2$$

But LCLS linewidth $\sim 8 \text{ eV}$!

Rabi-flopping on 1s - 2p resonance more feasible

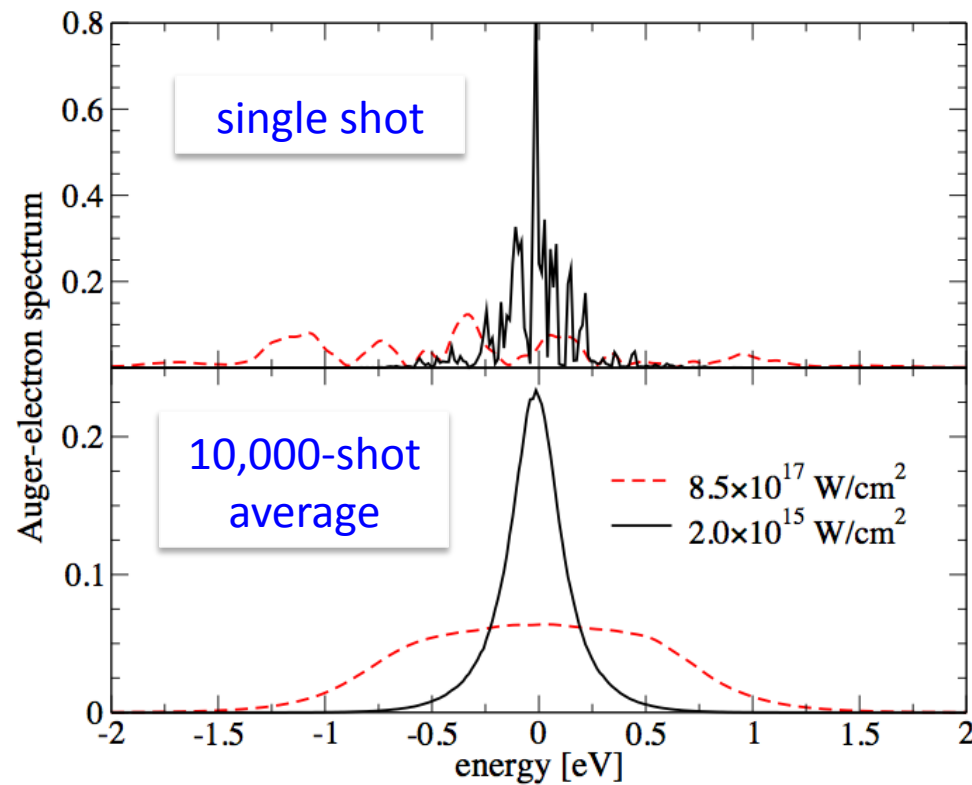


$$E_{x\text{-ray}} = 848.6 \text{ eV}$$

$$\sigma_{1s-2p} = 500\sigma_{2p-\infty} = 30 \sigma_{1s-3p}$$

Observe Auger yield when x-rays scanned over 1s - 2p resonance.
 Observe broadening at resonance to indicate Rabi flopping
 Theory: Rohringer & Santra PRA (2008).

Calculated “Resonant Auger effect at high x-ray intensity”



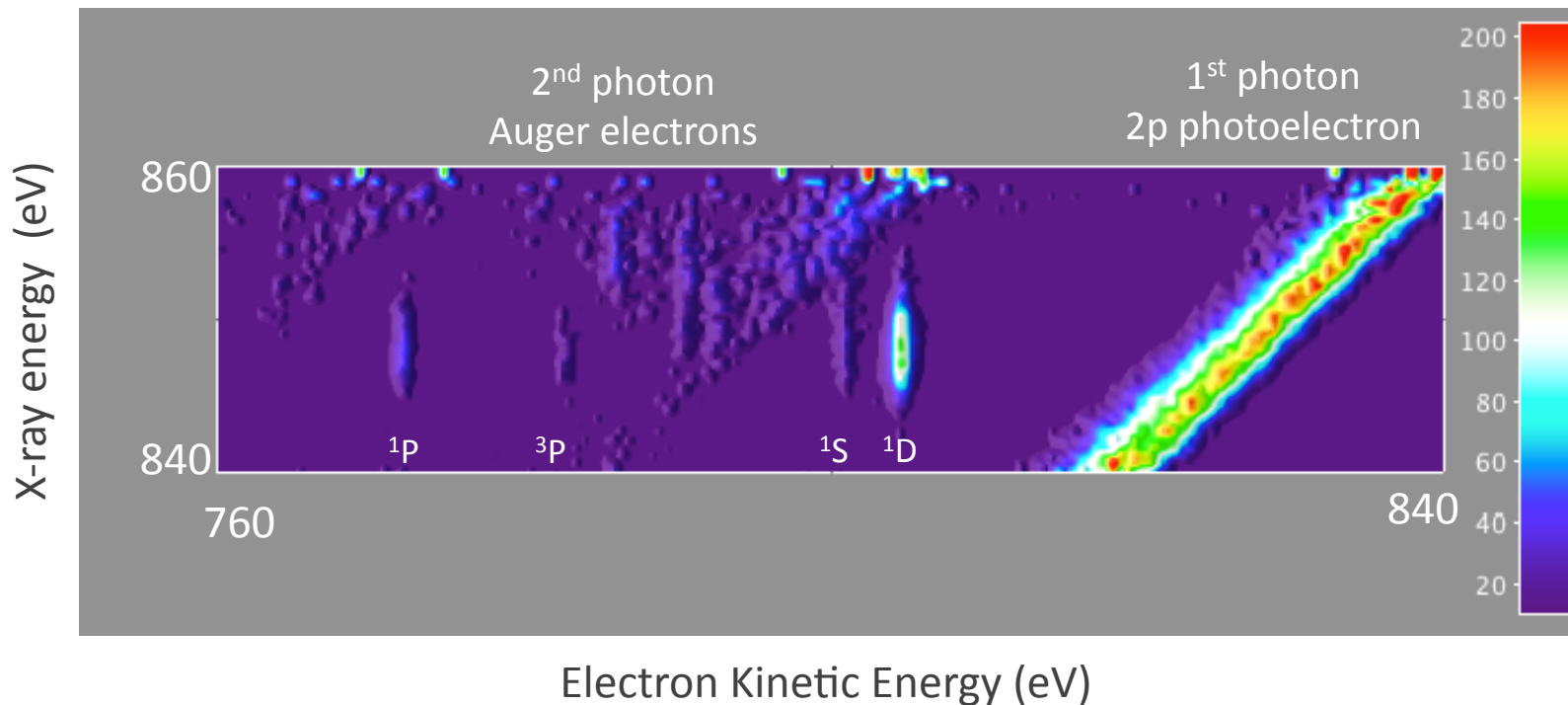
-> Look for Auger line broadening on resonance

N. Rohringer & R. Santra, PRA 77, 053404 (2008)



Electron spectra vs photon energy

using eTOF1 perpendicular to photon polarization to
suppress 2s photoelectrons (40 pC/bunch)



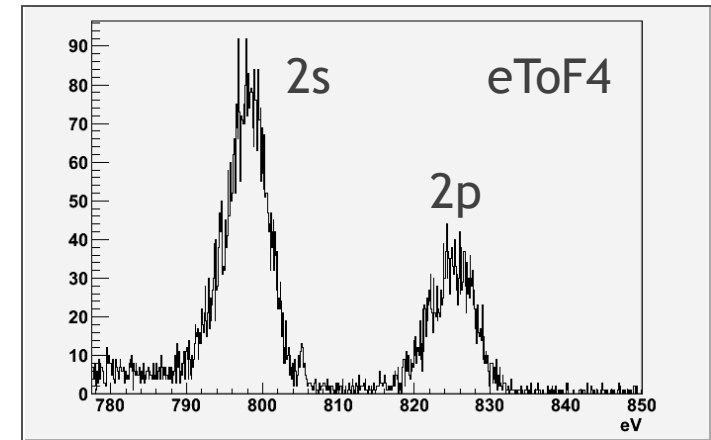
“Hidden” resonance revealed in single SASE pulse
Preliminary from Bertold Krässig



Characterization of X-ray SASE pulse

- Shot-to-shot photon energy jitter
 - Photoelectron energy spectra give x-ray energy centroid
 - Jitter derived from shot-to-shot GeV electron beam energy measurements

Conditions	FWHM photon energy jitter (eV)
40 pC (<10 fs) 850 eV 0.3 mJ 4500 A	4.25
250 pC (100 fs) 787 eV 1.5 mJ 2500 A	4.79
250 pC (100 fs) 769 eV 1.5 mJ 2500 A	5.24

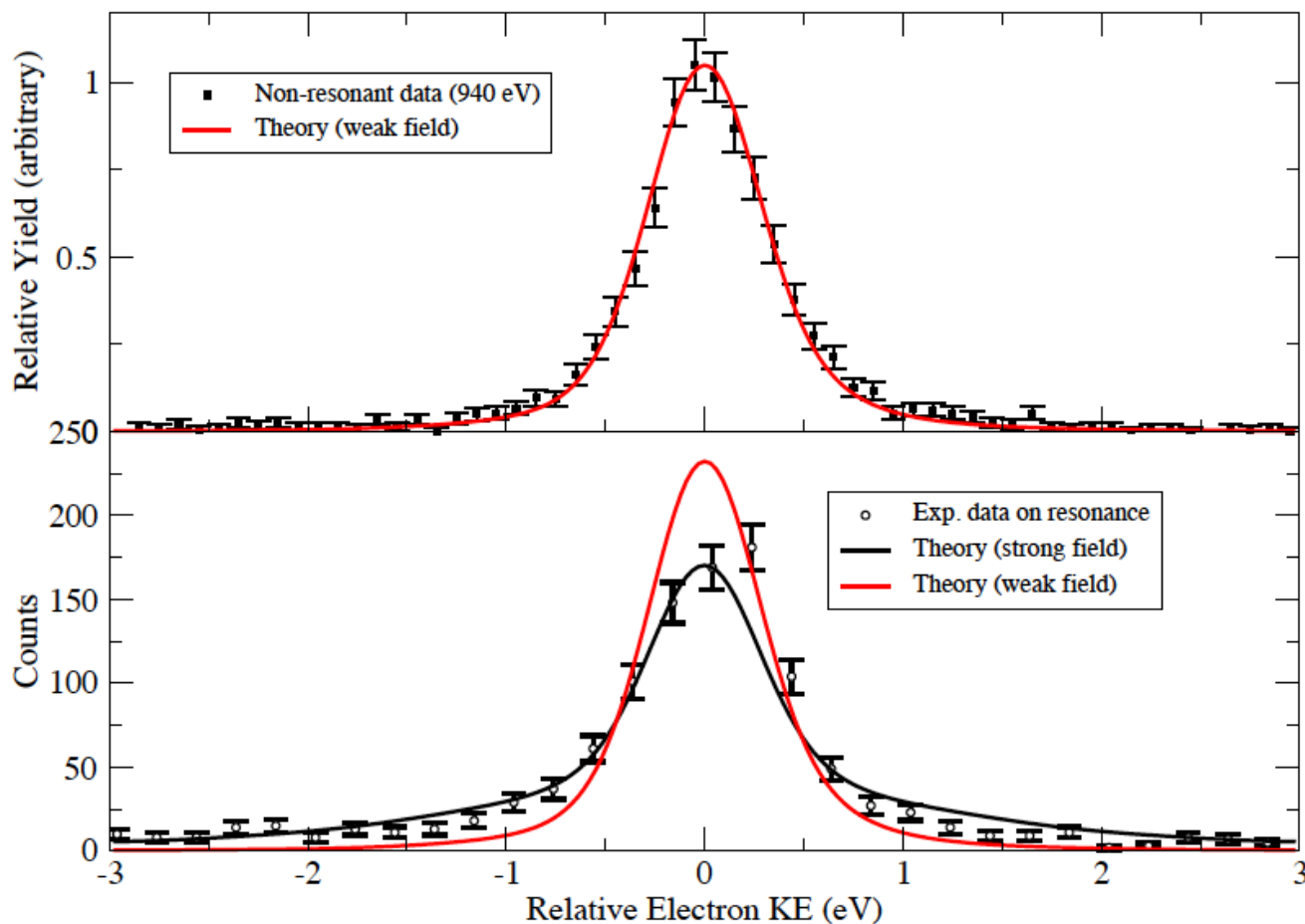


- Intrinsic x-ray bandwidth

Conditions	Intrinsic x-ray pulse bandwidth (from 2s photopeak) (eV) (FWHM)	Intrinsic x-ray pulse bandwidth (from 2p photopeak) (eV) (FWHM)	Average bandwidth (eV) (FWHM)	%
40 pC (<10 fs) 850 eV 0.3 mJ 4500 A	4.3	4.5	4.4	0.5 %
250 pC (100 fs) 787 eV 1.5 mJ 2500 A	7.1	7.8	7.45	0.9%
250 pC (100 fs) 769 eV 1.5 mJ 2500 A	7.7	7.8	7.77	1%



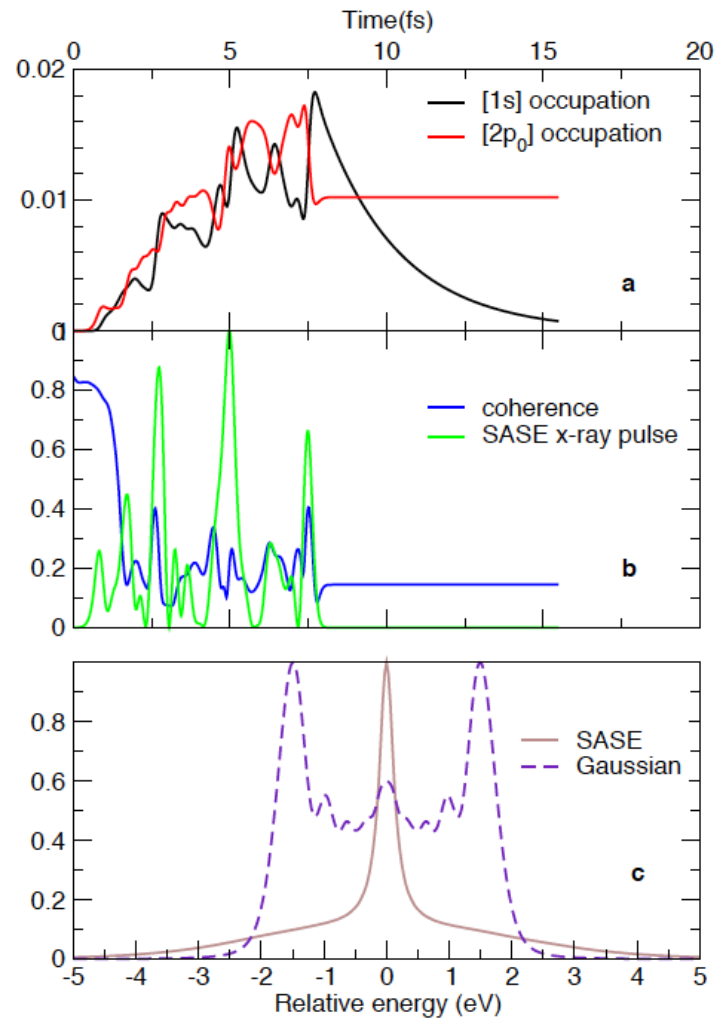
Is the 1D Auger line broadened on 1s-2p resonance?



X-ray parameters
0.3 mJ, 8.5 fs, $2 \mu\text{m}^2$
(20% transmission)

Theory from N. Rohringer and R. Santra
Preliminary Analysis – E. Kanter

Improved control of inner-shell dynamics with Gaussian pulse



$\sim 2 \times 10^{17} \text{ W/cm}^2$

0.3 mJ, 8.5 fs, $2 \mu\text{m}^2$

To be submitted
Kanter et al.
Calcs from N. Rohringer



Summary

- Insight into ultraintense x-ray interactions
 - six-photon, ten-electron stripping of neon ($\sim 10^{12}/\mu\text{m}^2$)
 - multiple photon absorption probability high when fluence $> 1/\sigma$
- Intensity-induced x-ray transparency – a general phenomena
 - transient x-ray transparency caused by formation of hollow atoms
 - hollow atoms $\sigma_{\text{scatt}}/\sigma_{\text{abs}}$ is increased
- Femtosecond time-scale atomic processes provide FEL diagnostics
- Straightforward rate equation calculations capture essential physics
- Hidden resonances are potential complications in using intense SASE pulses
- Intense x-rays can “control” inner-shell electron dynamics

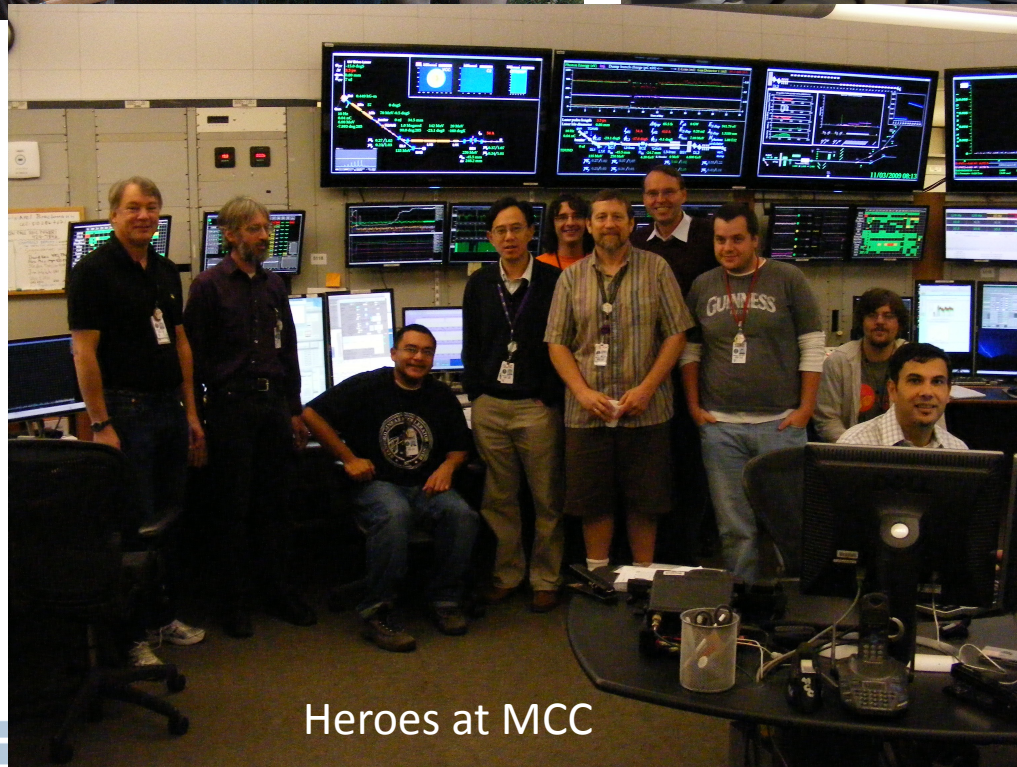




Argonne AMO group Oct 2009



Heroes at AMO Control



Heroes at MCC

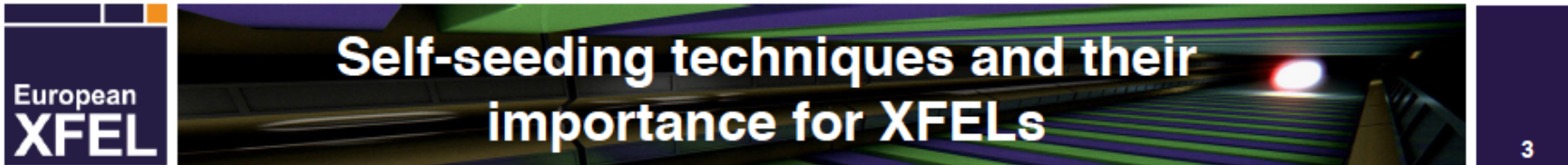




Outlook/Wider Impact/Final Thoughts



The chaotic truth about current XFELs



SASE pulses, baseline mode of operation: poor longitudinal coherence

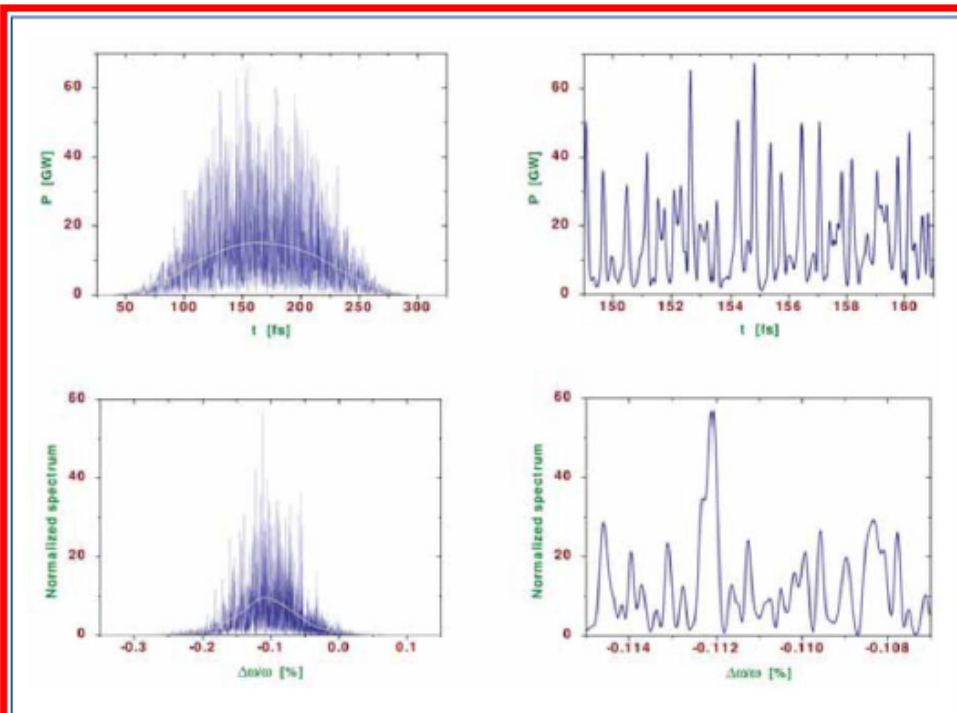


Figure 5.2.4 Temporal (top) and spectral (bottom) structure for 12.4 keV XFEL radiation from SASE 1. Smooth lines indicate averaged profiles. Right side plots show enlarged view of the left plots. The magnetic undulator length is 130 m.

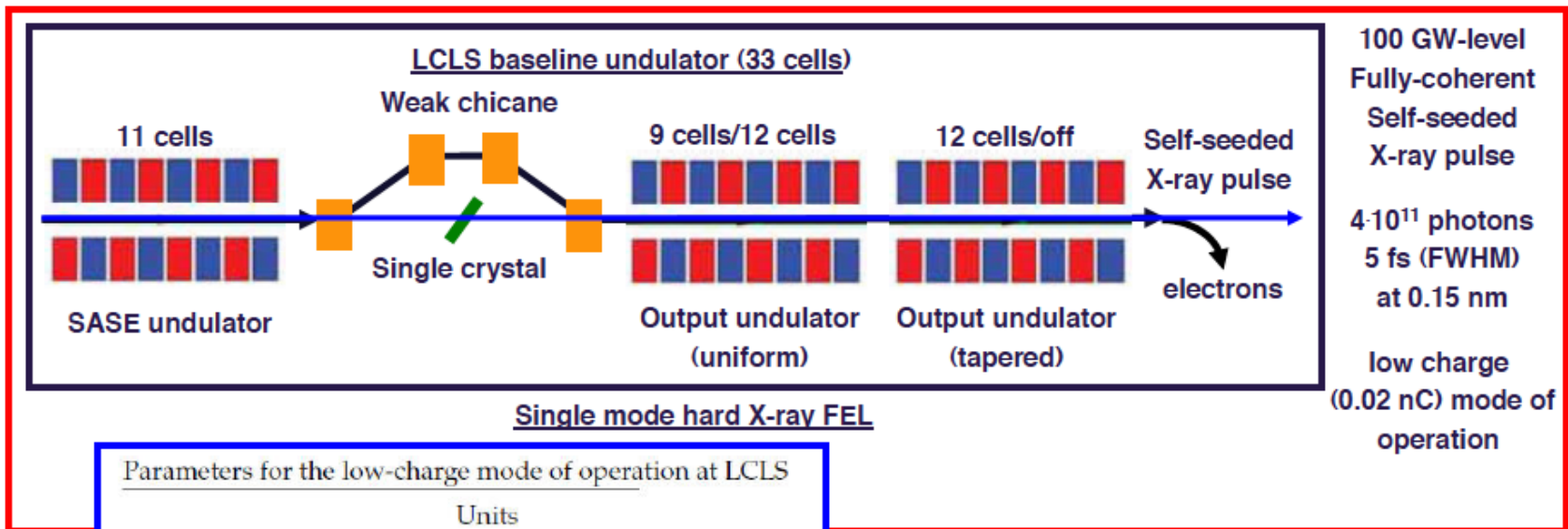
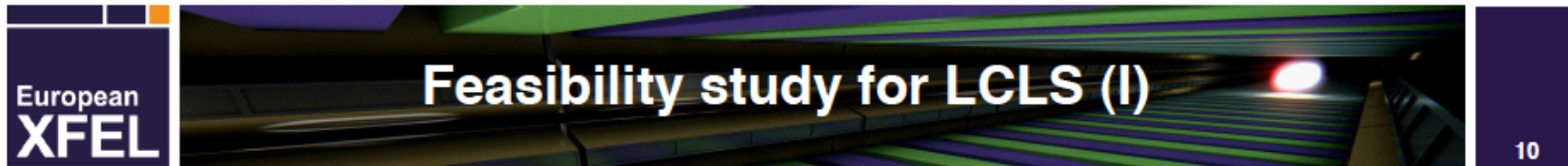
Source: The European XFEL TDR – DESY 2006-097 (2006)

$$\frac{\Delta\omega}{\omega} \sim 2\rho \sim 10^{-3}$$

$$\left(\frac{\Delta\omega}{\omega}\right)_{\text{spike}} \sim \frac{1}{\sigma_T \omega} \sim 10^{-5}$$

- Hundreds of longitudinal modes
- A lot of room for improvement
- Self-seeding schemes answer the call for increasing longitudinal coherence

How to obtain longitudinally coherent x-ray laser pulses?



Parameters for the low-charge mode of operation at LCLS

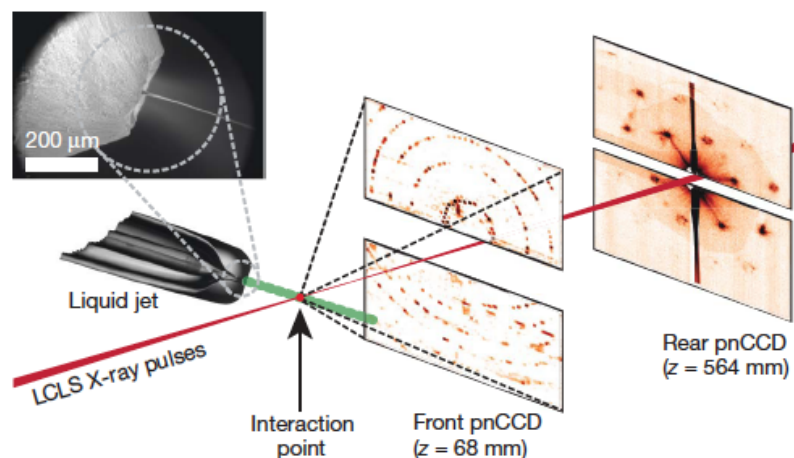
	Units	
Undulator period	mm	30
K parameter (rms)	-	2.466
Wavelength	nm	0.15
Energy	GeV	13.6
Charge	nC	0.02
Bunch length (rms)	μm	1
Normalized emittance	mm mrad	0.4
Energy spread	MeV	1.5

- 4m-long magnetic chicane
- $R_{56}=12\mu\text{m}$

Implementation in progress – SLAC/Argonne

Femtosecond X-ray protein nanocrystallography

Henry N. Chapman^{1,2}, Petra Fromme³, Anton Barty¹, Thomas A. White¹, Richard A. Kirian⁴, Andrew Aquila¹, Mark S. Hunter³, Joachim Schulz¹, Daniel P. DePonte¹, Uwe Weierstall⁴, R. Bruce Doak⁴, Filipe R. N. C. Maia⁵, Andrew V. Martin¹, Ilme Schlichting^{6,7}, Lukas Lomb⁷, Nicola Coppola^{1†}, Robert L. Shoeman⁷, Sascha W. Epp^{6,8}, Robert Hartmann⁹, Daniel Rolles^{6,7}, Artem Rudenko^{6,8}, Lutz Foucar^{6,7}, Nils Kimmel¹⁰, Georg Weidenspointner^{11,10}, Peter Holl⁹, Mengning Liang¹, Miriam Barthelmess¹², Carl Caleman¹, Sébastien Boutet¹³, Michael J. Bogan¹⁴, Jacek Krzywinski¹³, Christoph Bostedt¹³, Saša Bajt¹², Lars Gumprecht¹, Benedikt Rudek^{6,8}, Benjamin Erk^{6,8}, Carlo Schmidt^{6,8}, André Hömke^{6,8}, Christian Reich⁹, Daniel Pietschner¹⁰, Lothar Strüder^{6,10}, Günter Hauser¹⁰, Hubert Gorke¹⁵, Joachim Ullrich^{6,8}, Sven Herrmann¹⁰, Gerhard Schaller¹⁰, Florian Schopper¹⁰, Heike Soltau⁹, Kai-Uwe Kühnel⁸, Marc Messerschmidt¹³, John D. Bozek¹³, Stefan P. Hau-Riege¹⁶, Matthias Frank¹⁶, Christina Y. Hampton¹⁴, Raymond G. Sierra¹⁴, Dmitri Starodub¹⁴, Garth J. Williams¹³, Janos Hajdu⁵, Nicusor Timneanu⁵, M. Marvin Seibert^{5†}, Jakob Andreasson⁵, Andrea Rocker⁵, Olof Jönsson⁵, Martin Svenda⁵, Stephan Stern¹, Karol Nass², Robert Andritschke¹⁰, Claus-Dieter Schröter⁸, Faton Krasniqi^{6,7}, Mario Bott⁷, Kevin E. Schmidt⁴, Xiaoyu Wang⁴, Ingo Grotjohann³, James M. Holton¹⁷, Thomas R. M. Barends⁷, Richard Neutze¹⁸, Stefano Marchesini¹⁷, Raimund Fromme³, Sebastian Schorb¹⁹, Daniela Rupp¹⁹, Marcus Adolph¹⁹, Tais Gorkhover¹⁹, Inger Andersson²⁰, Helmut Hirsemann¹², Guillaume Potdevin¹², Heinz Graafsma¹², Björn Nilsson¹² & John C. H. Spence⁴

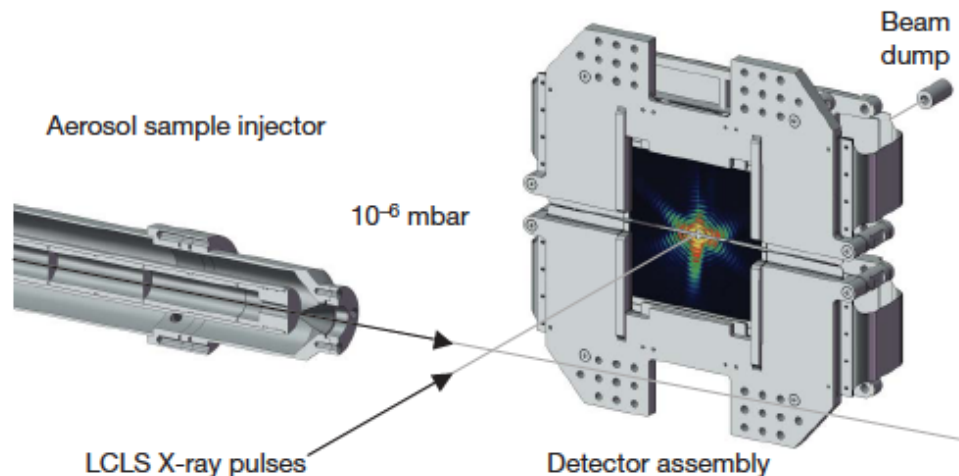


Photosystem I

membrane protein, 1 MDa
3,000,000 diffraction patterns
Crystals 200 nm - 2 μm
Resolution 8.5 Å
Match synchrotron structure
Dose: 700 MGy
Typical synch dose: 30 MGy

Single mimivirus particles intercepted and imaged with an X-ray laser

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Mimivirus

- Largest known virus – 0.75 μm
 - Does not crystallize
 - Too large for 3D cryoelectron microscopy
- Single Shot Scattering Pattern
32 nm resolution

Final thoughts

- XFELs open an entirely new intensity, wavelength regime for exploration
 - multiphoton effects are prominent
 - a complete contrast to any other x-ray source on earth
 - can create and probe exotic forms of matter
- XFEL research is in its infancy
 - Just learning how to use and control this source of radiation
 - LCLS is a superb opportunity to understand basics and explore novel ideas
 - Nanocrystallography appears to be an early success
- XFEL access is very limited
 - One beamline at LCLS currently
 - Beamlines at XFEL by 2015

