

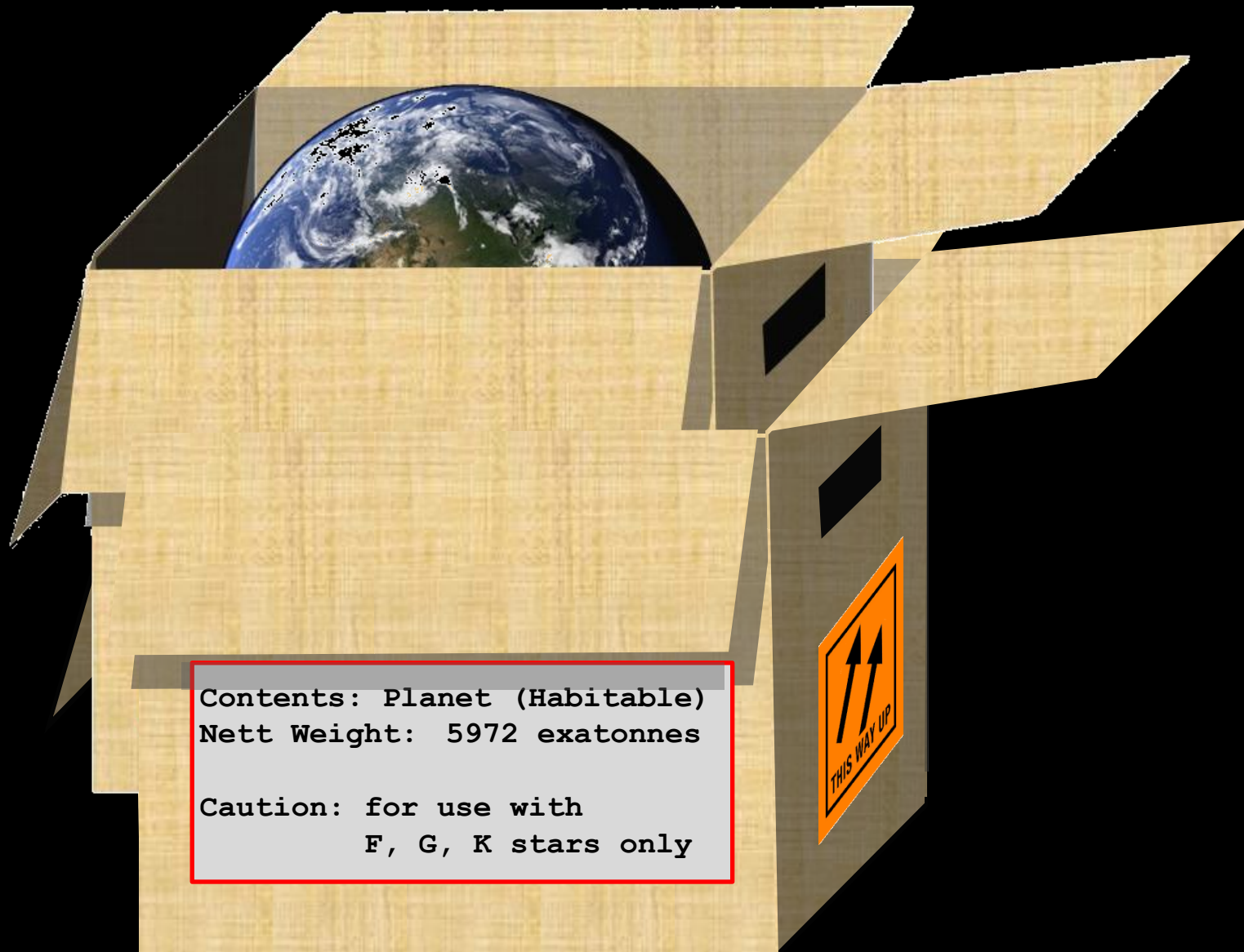
Extrasolar biosignatures: thinking outside the box

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Contents: Planet (Habitable)
Nett Weight: 5972 exatonnes

Caution: for use with
F, G, K stars only

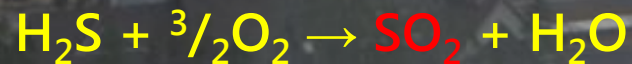
Non-terrestrial constraints on biosignatures

- Based on understanding of what life does:
 - What is life?
 - Why does it make volatiles
 - Why does it make colours
 - Only chemical 'laws' as assumptions
- Based on admitting we do not know the answers to that
 - Exhaustive list of possibilities: what does geology *not* make?

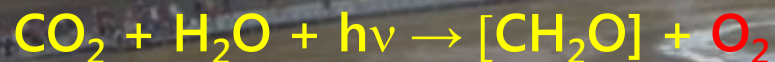


Biosignature gases

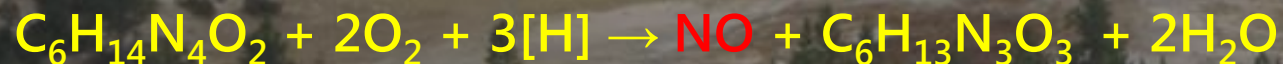
- Classified on why life makes them
 - Type I: byproduct of energy capture



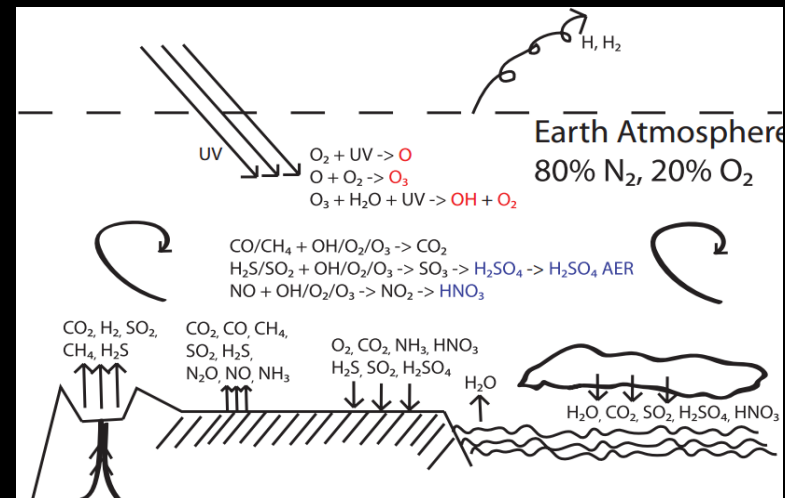
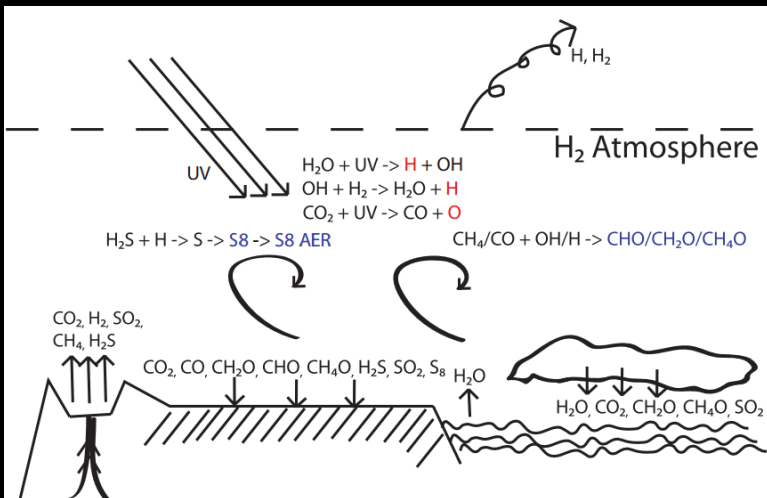
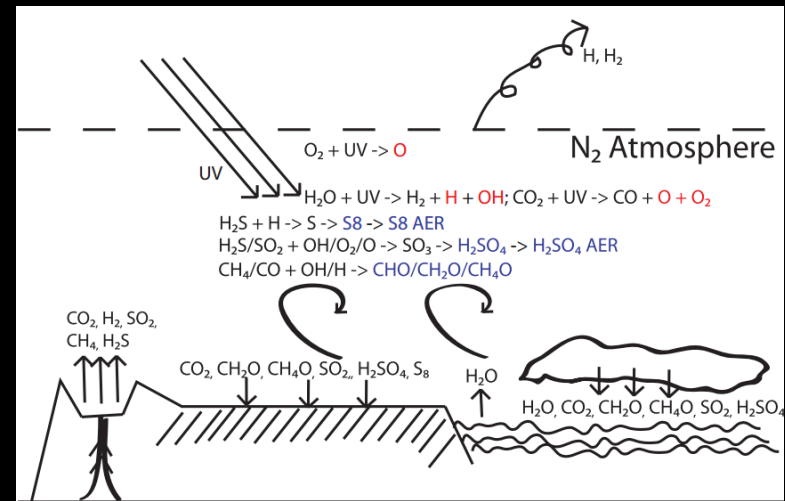
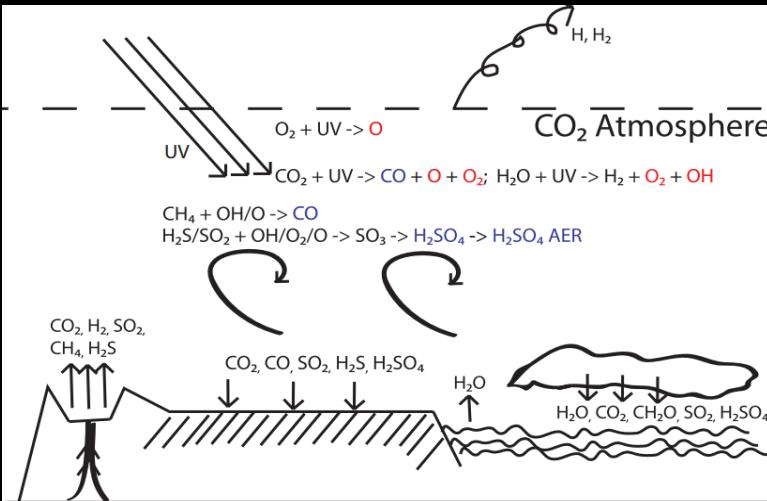
- Type II: byproduct of biomass capture



- Type III: no chemical “reason” at all



Type I prediction



Type I products on H₂-dominated worlds

Element	Reaction	ΔG° (kJ/mol)
Carbon	$\text{CO} + 3 \text{H}_2 \rightarrow \text{CH}_4 + \text{H}_2\text{O}$	-205.6
	$\text{CO}_2 + 4\text{H}_2 \rightarrow \text{CH}_4 + 2\text{H}_2\text{O}$	-194.5
	$\text{CO}_2 + \text{H}_2 \rightarrow \text{CO} + \text{H}_2\text{O}$	+11.5
	$\text{CO}_3^{2-} + 4\text{H}_2 \rightarrow \text{CH}_4 + \text{H}_2\text{O} + 2\text{OH}^-$	-129.42
Nitrogen	$\frac{1}{2} \text{N}_2 + 1\frac{1}{2} \text{H}_2 \rightarrow \text{NH}_3$	-62.61
	$\frac{1}{2} \text{N}_2 + \frac{1}{2} \text{H}_2 + \text{H}_2\text{O} \rightarrow \text{NH}_2\text{OH}$	+183.8
Phosphorus	$\text{H}_2 + \text{HPO}_4^{2-} \rightarrow \text{HPO}_3^{2-} + \text{H}_2\text{O}$	+27.2
	$\text{H}_2 + \text{HPO}_3^{2-} + \text{H}^+ \rightarrow \text{H}_2\text{PO}_2^- + \text{H}_2\text{O}$	+84.3
	$\frac{1}{2}\text{H}_2 + \text{H}_2\text{PO}_2^- + \text{H}^+ \rightarrow \text{P}_{(\text{s})} + 2\text{H}_2\text{O}$	+52.8
	$\text{P}_{(\text{s})} + 1\frac{1}{2}\text{H}_2 \rightarrow \text{PH}_3$	+5.4
	Overall $4 \text{H}_2 + \text{HPO}_4^{2-} + 2\text{H}^+ \rightarrow \text{PH}_3 + 4\text{H}_2\text{O}$	+169.8
Sulfur	$\text{SO}_4^{2-} + \text{H}_2 \rightarrow \text{SO}_3^{2-} + \text{H}_2\text{O}$	+12.45
	$\text{SO}_3^{2-} + 2\text{H}_2 + 2\text{H}^+ \rightarrow \text{S}_{(\text{s})} + 3\text{H}_2\text{O}$	-248.29
	$\text{S}_{(\text{s})} + \text{H}_2 \rightarrow \text{H}_2\text{S}$	-44.81
	Overall $\text{SO}_4^{2-} + 2\text{H}^+ + 4\text{H}_2 \rightarrow \text{H}_2\text{S} + 4\text{H}_2\text{O}$	-280.8
Iron	$\frac{1}{2} \text{H}_2 + \text{Fe}^{3+} + \text{OH}^- \rightarrow \text{Fe}^{2+} + \text{H}_2\text{O}$	-125.8
	$\text{H}_2 + \text{Fe}^{2+} + 2\text{OH}^- \rightarrow \text{Fe}_{(\text{s})} + 2\text{H}_2\text{O}$	-6.1
Manganese	$\text{Mn}^{3+} + \frac{1}{2} \text{H}_2 + \text{OH}^- \rightarrow \text{Mn}^{2+} + \text{H}_2\text{O}$	-273.3
	$\text{H}_2 + \text{Mn}^{2+} + 2\text{OH}^- \rightarrow \text{Mn}_{(\text{s})} + 2\text{H}_2\text{O}$	-24.9
Silicon	$2\text{H}_2 + \text{H}_4\text{SiO}_4_{(\text{s})} \rightarrow \text{Si}_{(\text{s})} + 4\text{H}_2\text{O}$	+384.5
	$2\text{H}_{2(\text{g})} + \text{SiO}_{2(\text{s})} \rightarrow \text{Si}_{(\text{s})} + 2\text{H}_2\text{O}$	+382.1
	$\text{Si}_{(\text{s})} + 2\text{H}_{2(\text{g})} \rightarrow \text{SiH}_{4(\text{g})}$	+56.9
Aluminium	$3\text{H}_{2(\text{g})} + \text{Al}_2\text{O}_{3(\text{s})} \rightarrow 2\text{Al}_{(\text{s})} + 3 \text{H}_2\text{O}$	+871.0
Copper	$\text{Cu}^{2+} + \frac{1}{2}\text{H}_2 \rightarrow \text{Cu}^+ + \text{H}^+$	-19.4
	$\text{Cu}^+ + \frac{1}{2} \text{H}_2 \rightarrow \text{Cu}_{(\text{s})} + \text{H}^+$	-57.8
Vanadium	$\text{H}_2\text{VO}_4^- + 2\text{H}^+ + \frac{1}{2}\text{H}_2 \rightarrow \text{HVO}_2^+ + 2\text{H}_2\text{O}$	-113.8
	$\text{HVO}_2^+ + \frac{1}{2}\text{H}_2 \rightarrow \text{VO}^+ + \text{H}_2\text{O}$	-243.5
	$\text{VO}^+ + \frac{1}{2} \text{H}_2 \rightarrow \text{VOH}^+$	17.5
	$\text{VOH}^+ + \text{H}_2 \rightarrow \text{V}_{(\text{s})} + \text{H}^+ + \text{H}_2\text{O}$	122.7

+ Na, K, Ca, Mg, etc – too electropositive

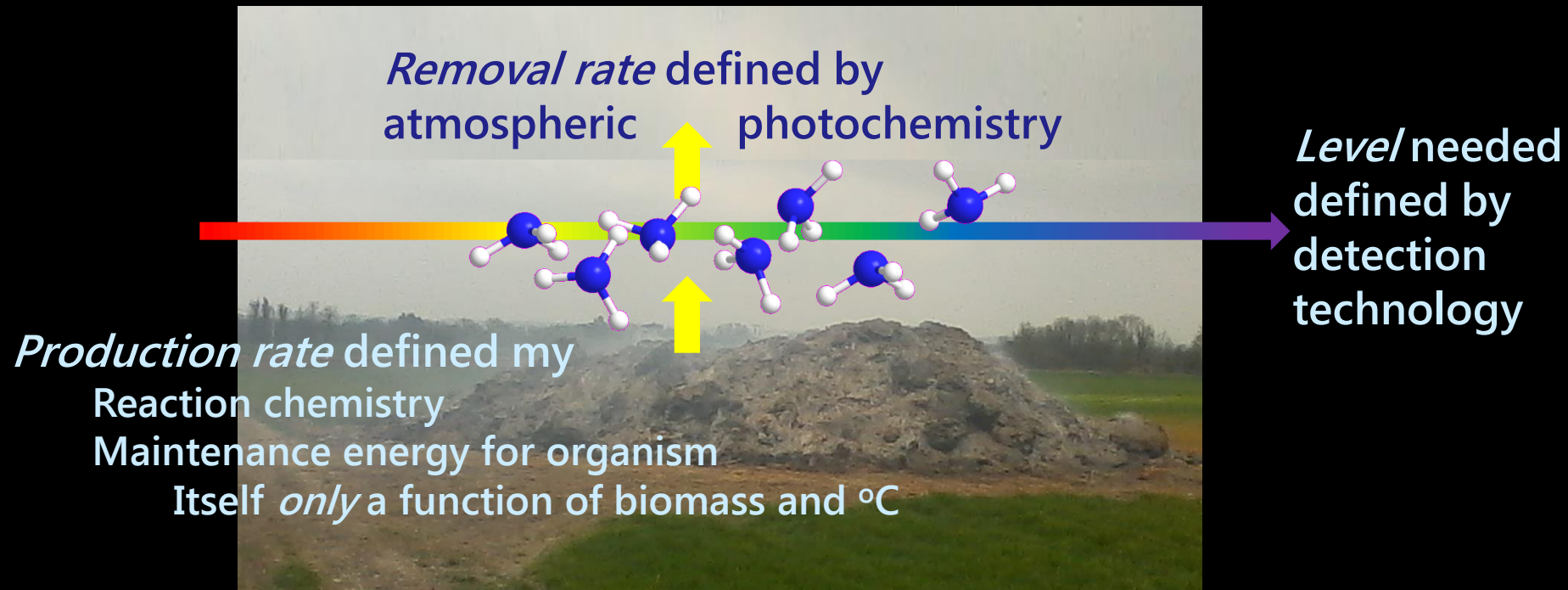
- Plausible substrates
- Volatile products
- 25°C temp

–CH₄, H₂S, H₂O, **NH₃**



Biomass calculation

- Estimate biomass:



- Assumptions:

- Only one Type I reaction (implausible)
- No gas uptake (implausible)

Estimate is minimum



Live biomass densities



	Sun-like star	Active M star	Quiet M star
Thermal emission	$4.0 \times 10^{-4} \text{ (g/m}^2\text{)}$	$8.0 \times 10^{-6} \text{ (g/m}^2\text{)}$	$9.5 \times 10^{-6} \text{ (g/m}^2\text{)}$
Transmission	-	$1.1 \text{ (g/m}^2\text{)}$	$1.8 \times 10^{-9} \text{ (g/m}^2\text{)}$



10 - 50g/m²



50 - 1000 g/m²



0.01g/m²



Seager et al 2013b

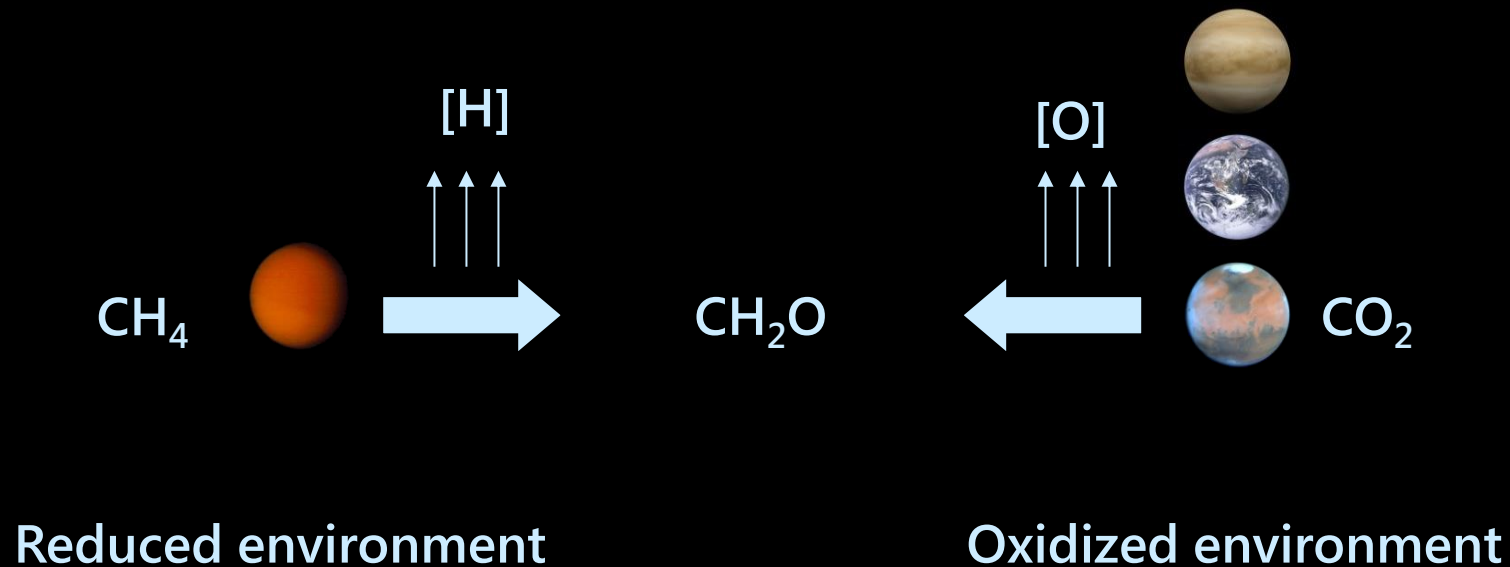
Problems

- False positives
 - Geochemistry has same chemicals to work with
 - E.G. methane on Mars?
- False negatives?
 - $\text{N}_2 + \text{O}_2 + \text{H}_2\text{O} \rightarrow \text{HNO}_3$
 - abundant source gas
 - BEQ



Gases: Type II

- 'Problem' for life
- To access large chemical space, need molecules R_f 0.5 – 0.9
- Environmental carbon in *dense* bodies



Redox signature in presence of H₂

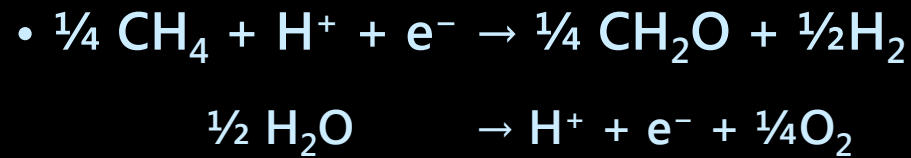
Outgassing	Other process	Atmospheric carbon
CO ₂ / CO	-	H ₂ + 10:1 CO ₂ :CH ₄
CO ₂ / CH ₄	-	H ₂ + 50:50 CO ₂ :CH ₄
CO ₂ / CH ₄	CO ₂ + H ₂ → CH ₄	H ₂ + 1:99 CO ₂ :CH ₄

- Type II signature
- CH₄ + H₂O + energy → [CH₂O]_n + [H]
- Most plausibly
- CH₄ + H₂O + energy → [CH₂O]_n + H₂

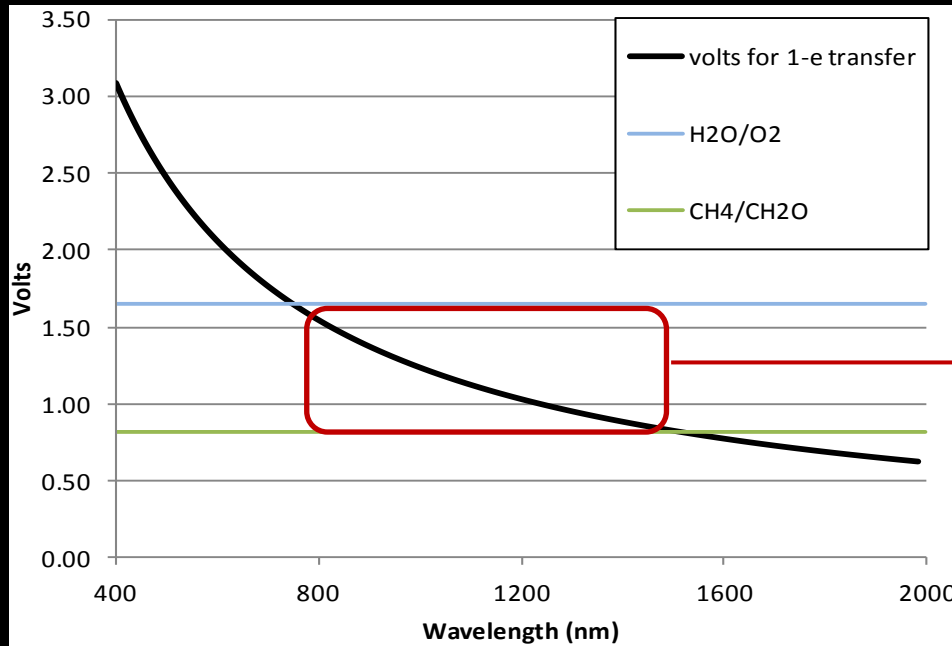
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	CO ₂ + H ₂ → CO + H ₂ O	+11.5
	CO ₃ ²⁻ + 4 H ₂ → CH ₄ + H ₂ O + 2 OH ⁻	-129.42
Nitrogen	½ N ₂ + 1 ½ H ₂ → NH ₃	-62.61
	½ N ₂ + ½ H ₂ + H ₂ O → NH ₃ OH	+183.8
Phosphorus	H ₂ + HPO ₄ ²⁻ → HPO ₃ ²⁻ + H ₂ O	+27.2
	H ₂ + HPO ₄ ²⁻ + H ⁺ → H ₂ PO ₂ ⁻ + H ₂ O	+84.3
	½ H ₂ + H ₂ PO ₂ ⁻ + H ⁺ → P _(s) + 2 H ₂ O	+52.8
	P _(s) + 1 ½ H ₂ → PH ₃	+5.4
Overall 4 H ₂ + HPO ₄ ²⁻ + 2 H ⁺ → PH ₃ + 4 H ₂ O		+169.8
Sulfur	SO ₄ ²⁻ + H ₂ → SO ₃ ²⁻ + H ₂ O	+12.45
	SO ₃ ²⁻ + 2 H ₂ + 2 H ⁺ → S _(s) + 3 H ₂ O	-248.29
	S _(s) + H ₂ → H ₂ S	-44.81
Overall SO ₄ ²⁻ + 2 H ⁺ + 4 H ₂ → H ₂ S + 4 H ₂ O		-280.8
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	H ₂ + Mn ²⁺ + 2 OH ⁻ → Mn _(s) + 2 H ₂ O	-24.9
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	2 H _{2(g)} + SiO _{2(g)} → Si _(s) + 2 H ₂ O	+382.1
	Si _(s) + 2 H _{2(g)} → SiH _{4(g)}	+56.9
Aluminium	3 H _{2(g)} + Al ₂ O _{3(s)} → 2 Al _(s) + 3 H ₂ O	+871.0
Copper	Cu ²⁺ + ½ H ₂ → Cu ⁺ + H ⁺	-19.4
	Cu ⁺ + ½ H ₂ → Cu _(s) + H ⁺	-57.8
Vanadium	H ₂ VO ₄ ⁻ + 2 H ⁺ + ½ H ₂ → HVO ₂ ⁺ + 2 H ₂ O	-113.8
	HVO ₂ ⁺ + ½ H ₂ → VO ⁺ + H ₂ O	-243.5
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VOH ⁺ + H ₂ → V _(s) + H ⁺ + H ₂ O		122.7



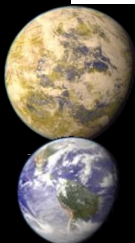
Photon energy



$$\left. \begin{aligned} E^\circ &= 0.63 \text{ V} \\ E^\circ &= 1.23 \text{ V} \end{aligned} \right\} + \text{overvoltage}$$



Absorption in NIR
→ no 'Red Edge'?

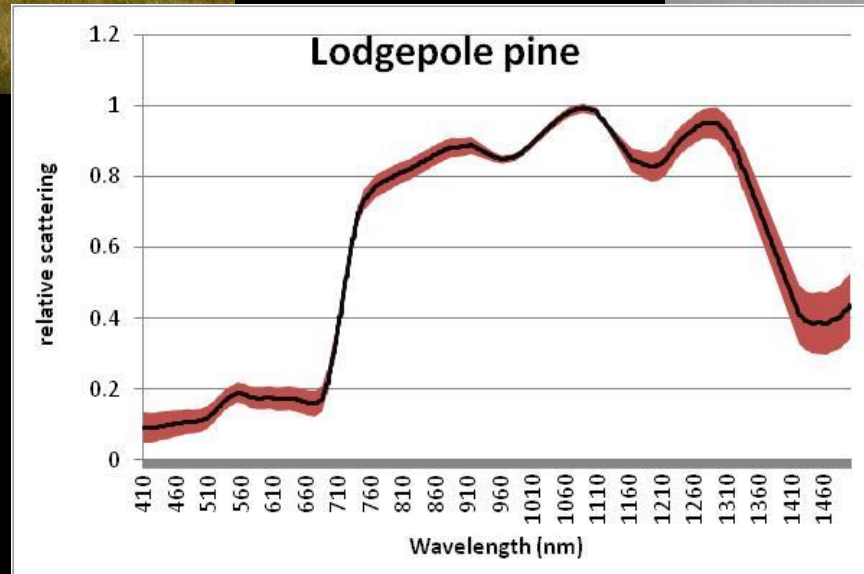


Problems

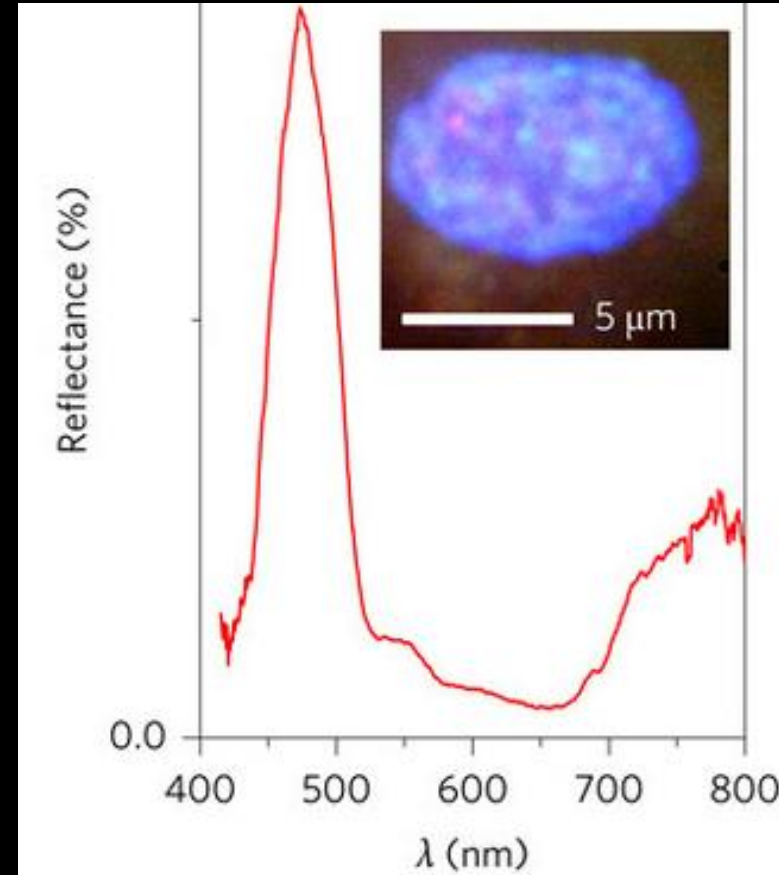
- False positives
 - Less problematic – thermodynamically implausible
- Evolutionary negatives
 - Earth pre-GEO
 - Earth pre-Cambrian?
- Indetectability in reduced environments



Surface color: The Red Edge



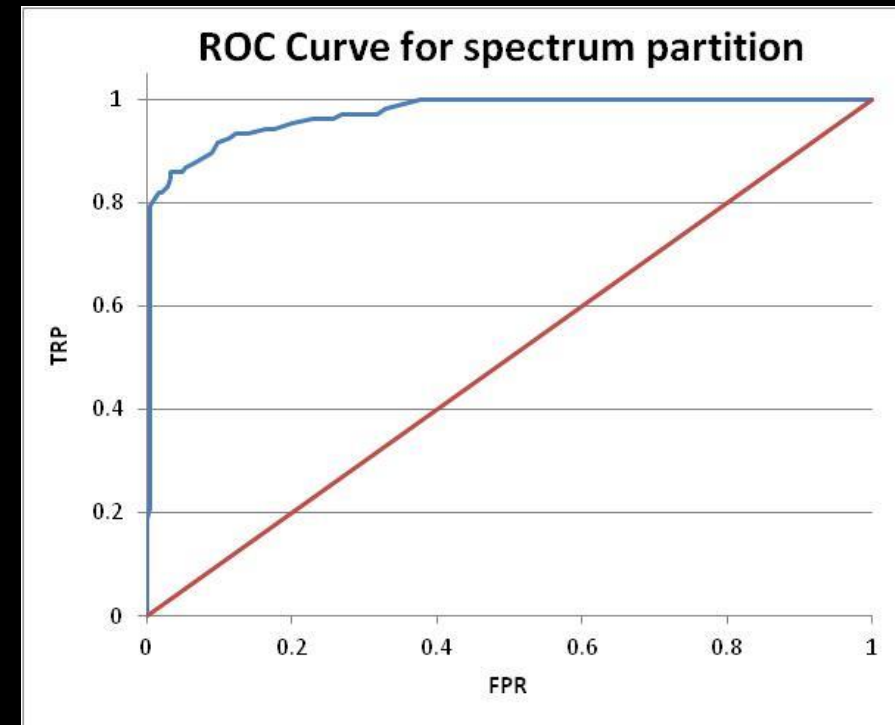
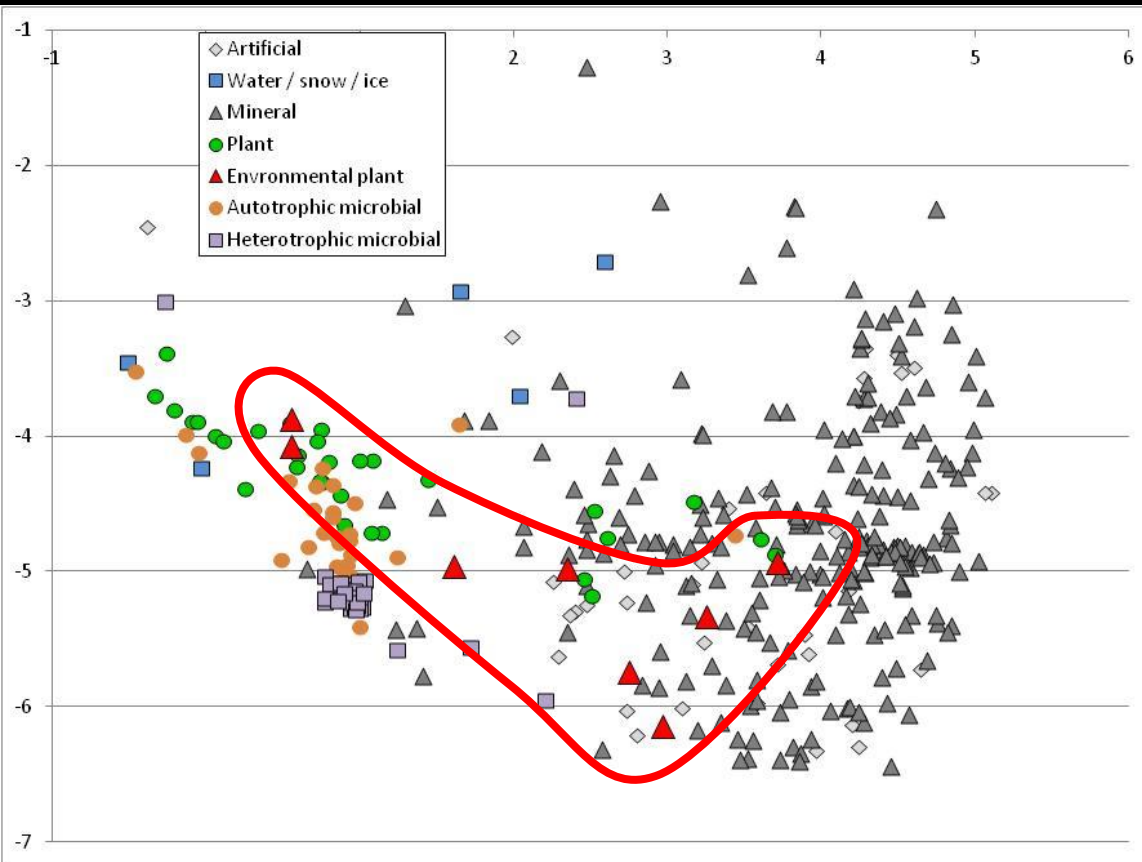
Chlorophyll 'Red Edge' is mutable



- Photonic multilayer structure of *Begonia* chloroplasts enhances photosynthetic efficiency. Jacobs et al 2016

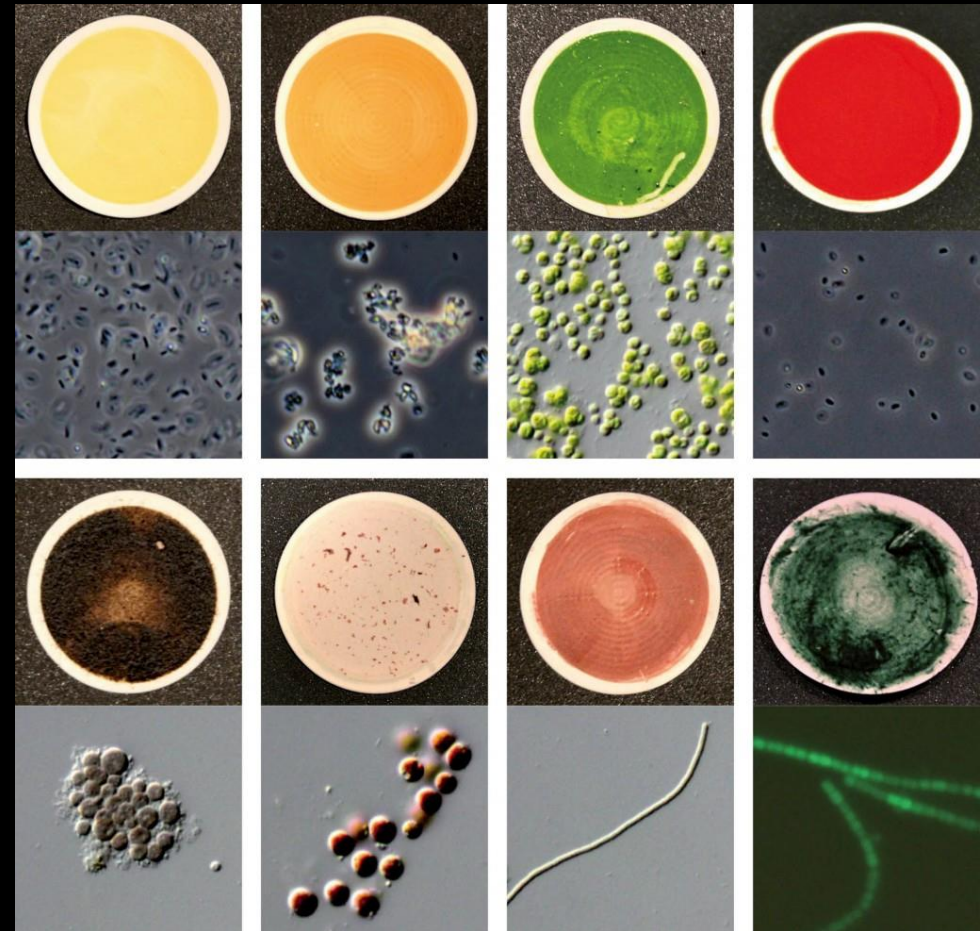


Spectral classification (PCA)



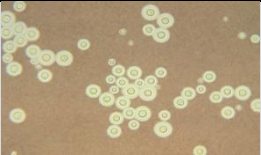
Colour of life

- Can you derive colour from function?
- Light harvesting
- UV/light protection
 - Melanin
- Visual signalling
 - Anything!
- Accidental colour
 - E.g. polyphenols



Energy capture $\Rightarrow$ pigments

- At least 4 independent light energy capture mechanisms
- At least 4 CO₂ assimilation pathways
- Only one winner ...

Pigment system	Chlorophyll	Bacteriorhodopsins	Aphid carotenes	Melanized fungi
'Red Edge'	700 – 750	600 – 700	~550	NONE
				
Calvin cycle	✓	✗	✗	✗
rTCA cycle	✗	✗	✗	✗
3-OH-propionate	✗	✗	✗	✗
Reductive Acetyl-CoA	✗	✗	✗	✗

Mostly from *The evolution of photosynthesis...again?* Rothschild 2008



Gases: Type III

- Range of terrestrial Type III gases.
- Assume maximum terrestrial emission rates / gram biomass

Biomass needed to maintain detectable levels of Type III gases in atmosphere of 1bar H₂-dominated atmosphere planet in 'habitable zone' around different stars (g/m²)

Compound	Thermal emission			Transmission	
	Sun-like	Active M-star	Quiet M-star	Active M-star	Quiet M-star
CH ₃ Cl	2800	77	0.013	860	0.014
DMS	190	82	0.0001	260	0.00036
CS ₂	5.5×10^7	2.3×10^7	37	1.5×10^7	24
OCS	1.3×10^5	5500	0.67	9.9×10^4	12



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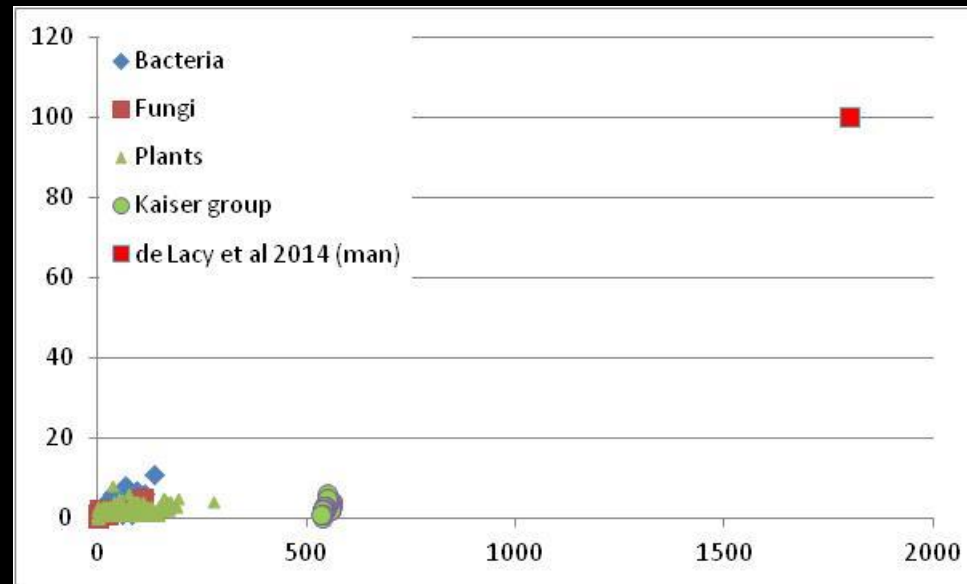
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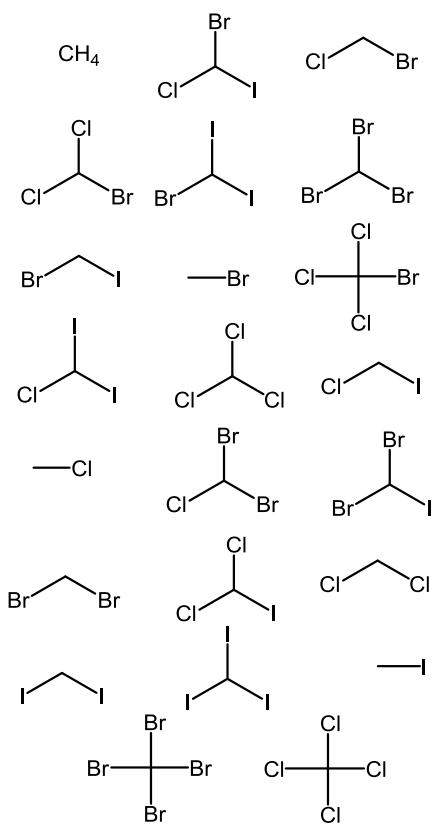
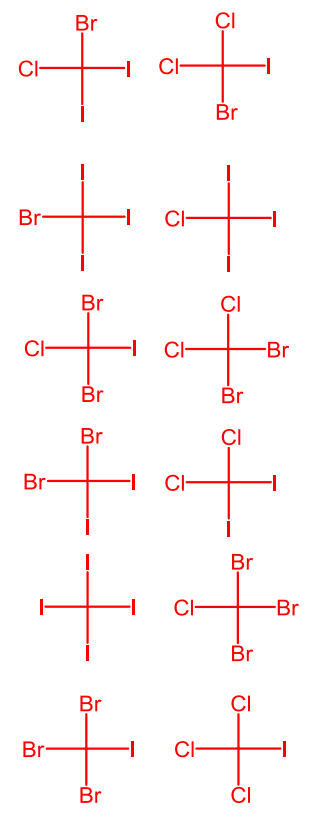
But ...

- What is collected?

- What is known?



EG halomethanes

Made by life	Not known to be made by life
	

- There are 34 possible halomethanes (excluding F). 22 are produced by Earth's life. 12 are not known to be produced. Why?
- Why not CFCs?
- Terrestrial life rarely handles F
- 2.4Gya rarely handled O



All Small Molecules project

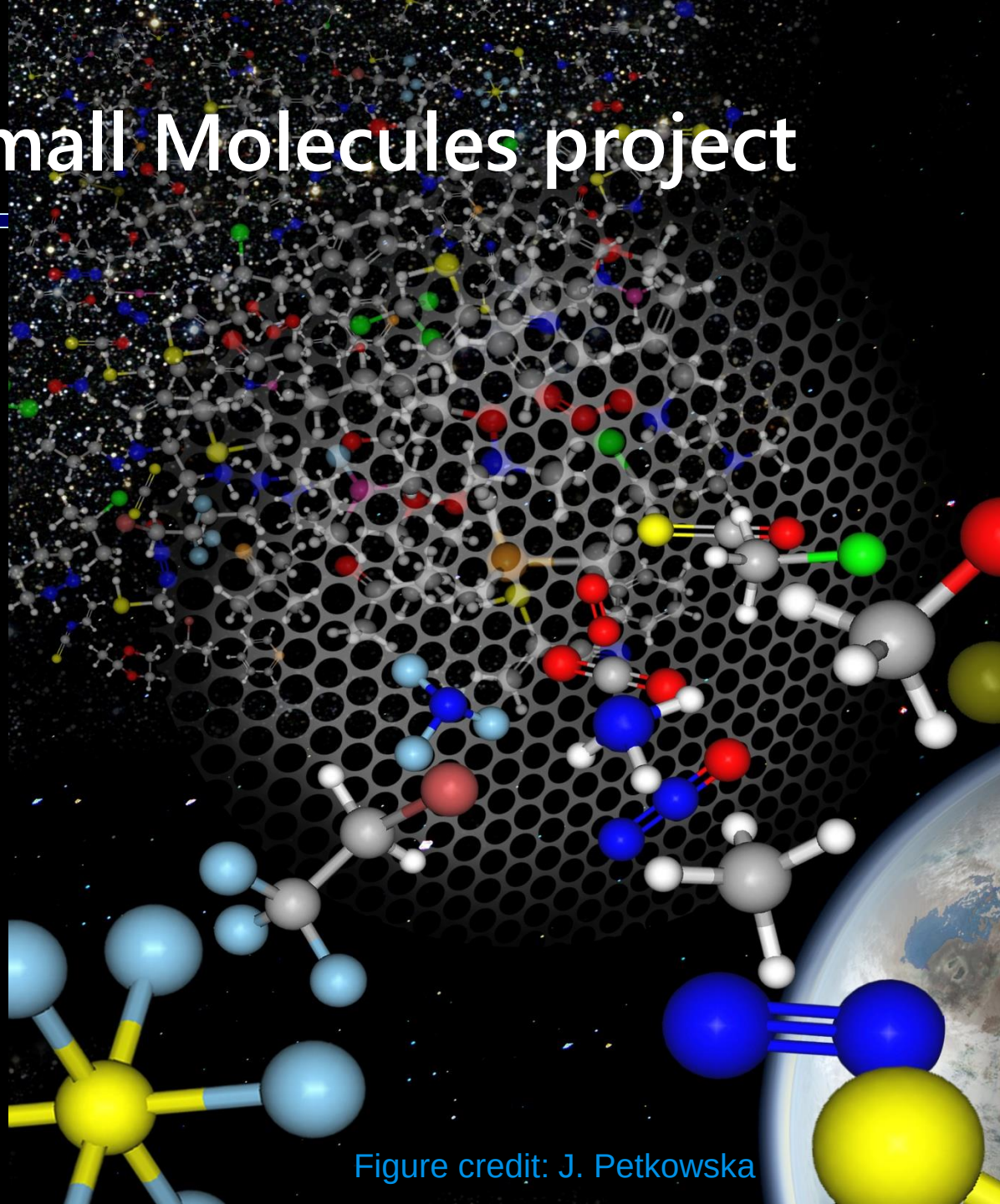
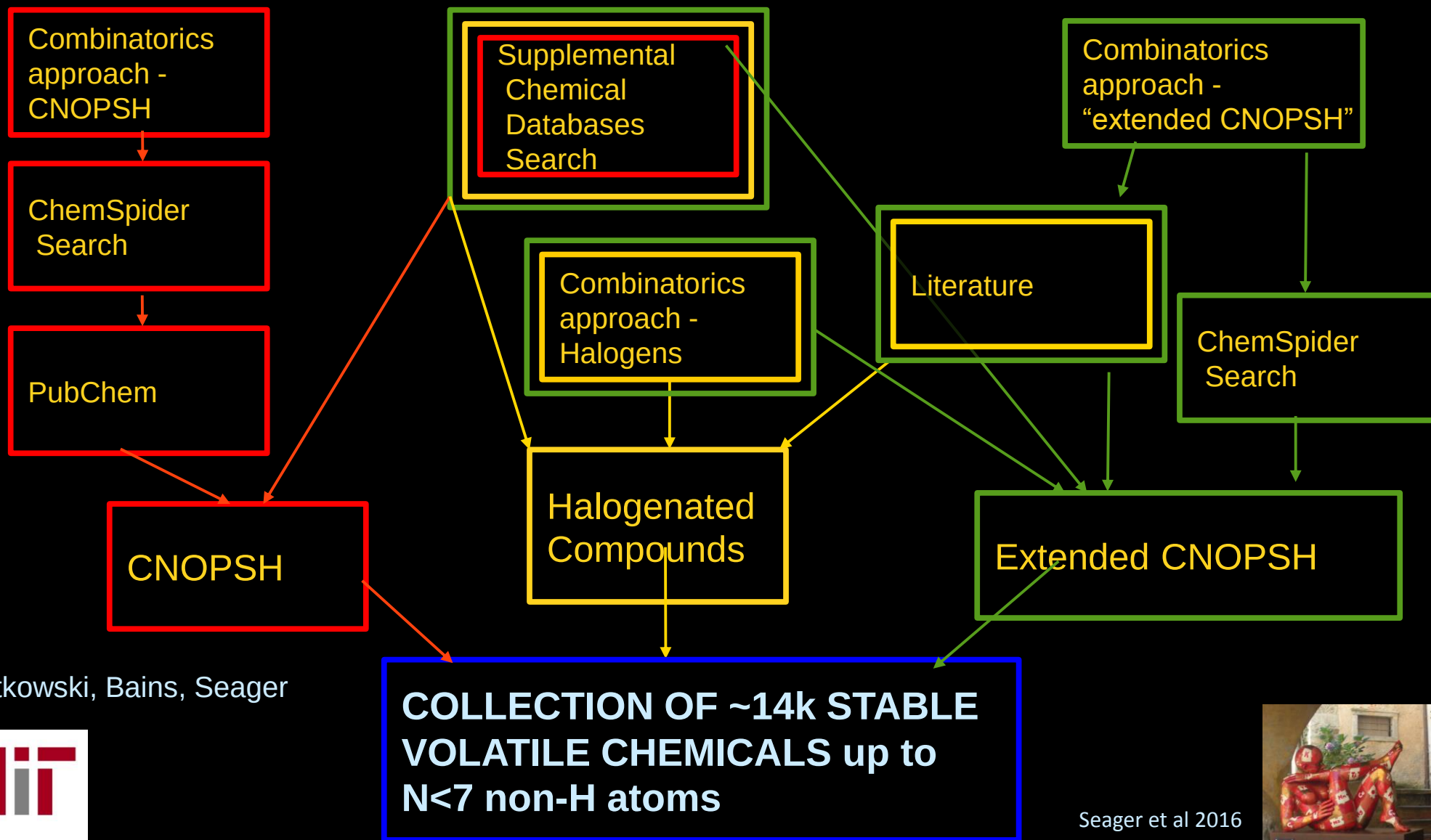


Figure credit: J. Petkowska



PBS database construction



Petkowski, Bains, Seager



Seager et al 2016

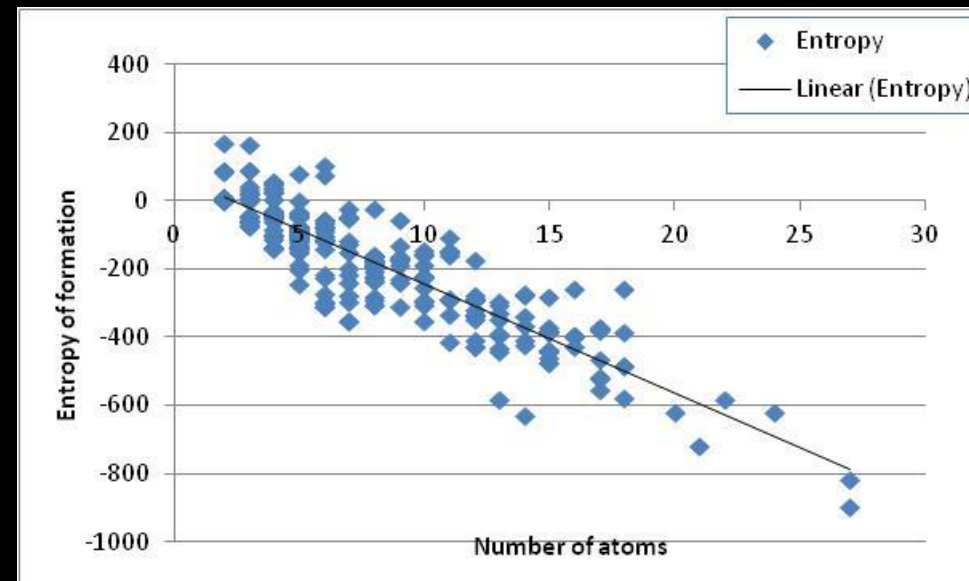


IUPAC name	Molecular Formula	Molecular Weight (Da)	SMILES	Boiling point (°C)	Produced by life Y/N	Reference for production by life
(...)						
N,N-Dimethylmethanamine	C3H9N	59.11	CN(C)C	4(e)	Y	Kite, G.C. and Hettterscheid, W.L.A. 1997
Buta-2,3-dien-1-amine	C4H7N	69.1	NCC=C=C	88(p)	N	
3-Buten-2-amine	C4H9N	71.12	C=CC(C)N	73(p)	N	
2-Oxiranecarbaldehyde	C3H4O2	72.06	O=CC1OC1	96(p)	N	
N-Ethylethanamine	C4H11N	73.14	CCNCC	55(e)	Y	Smith, B.N. and Meeuse, B.J.D. 1966
Hydrazinecarboxamide	CH5N3O	75.07	NNC(N)=O	210(p)	Y	van Poucke, C. et al. 2011
Amino(hydroxyamino)methanol	CH6N2O2	78.07	OC(N)NO	221(p)	N	
Cyclopropylmethanethiol	C4H8S	88.17	SCC1CC1	111(p)	N	
1,3-Dithietane 1-oxide	C2H4OS2	108.18	O=S1CSC1	198(p)	N	
Methyl hydrogen carbonotrithioate	C2H4S3	124.25	S=C(S)SC	200(p)	N	
(...)						
Stibine	H3Sb	124.78	[SbH3]	-18(e)	Y	Michalke, K. et al. 2000
3-(Methylselanyl)-1-propene	C4H8Se	135.07	C[Se]CC=C	113(p)	N	
Tetrahydroselenophene	C4H8Se	135.07	[Se]1CCCC1	128(p)	N	
2-(Methylselanyl)-1-propene	C4H8Se	135.07	[Se](C=C)C	107(p)	N	
N,N-Dimethylselenoformamide	C3H7NSe	136.05	[Se]=CN(C)C	120(p)	N	
Se-Methyl ethaneselenoate	C3H6OSe	137.04	O=C([Se]C)C	137(p)	N	
Methylstibine	CH5Sb	138.81	C[SbH2]	41(e)	Y	Wehmeier, S. et al. 2005
[(Methylselanyl)sulfanyl]methane	C2H6SSe	141.09	C[Se]SC	135(p)	N	
3-(Methyltellanyl)-1-propene	C4H8Te	183.71	[Te](CC=C)C	111(p)	N	
Monomethyl bismuth hydride	CH5Bi	226.03	C[BiH2]	80(p)	Y	Meyer, J. et al. 2008
(...)						
Bromotrichloromethane	CBrCl3	198.27	C(Cl)(Cl)(Cl)Br	105(e)	N	
Bromodiiodomethane	CHBrI2	346.73	IC(I)Br	221.5(p)	Y	Gribble, G. 2003
Chlorotribromomethane	CBBrCl	287.18	C(Cl)(Br)(Br)Br	84(e)	Y	Balschmiter, K. 2003
Trichloroacetonitrile	C2Cl3N	144.38	N#CC(Cl)(Cl)C	84(e)	Y	
1-Fluoropropane	C3H7F	62.09	CCCCF	34(e)	N	
Tetrafluoromethane	CF4	88	C(F)(F)(F)F	-150(e)	N	
Cyanic chloride	CNCl	61.47	C(#N)Cl	13(e)	N	
2-Bromo-1,1-dichloroethene	C2HBrCl2	175.84	C=C(Cl)Cl)Br	109(p)	Y	Nightingale, P.D. et al. 1995
Sulfur hexafluoride	SF6	146.06	FS(F)(F)(F)F	-64(e)	N	
Trichloroamine	NCI3	120.37	N(Cl)(Cl)Cl	71(e)	N	

Stuff that (non-industrial) terrestrial life is *extremely unlikely* to make

What next

- NIR detectability
 - Known geochemical production
 - Thermodynamics
 - Given range of modelled surface geochemistry
 - Local? Globally averaged?
 - E.g . Entropy of formation as guide
- $$\Delta G = \Delta H - T.\Delta S$$
- Develop 'geochemistry plausibility index'



280 small volatile molecules, CHONSP + B, Si, As, Ge



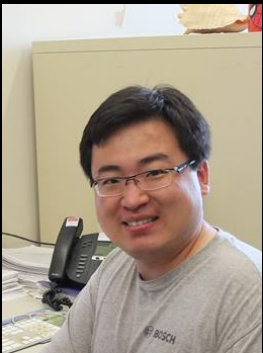
No man is an island ...



Sara Seager



Dirk Schulze-Makuch



Renyu Hu



Andras Zsom



Janusz Petkowski



Extrasolar biosignatures: thinking outside the box

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