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# ARCHIVING GROUND-BASED DATA: PERSPECTIVE FROM SPACE

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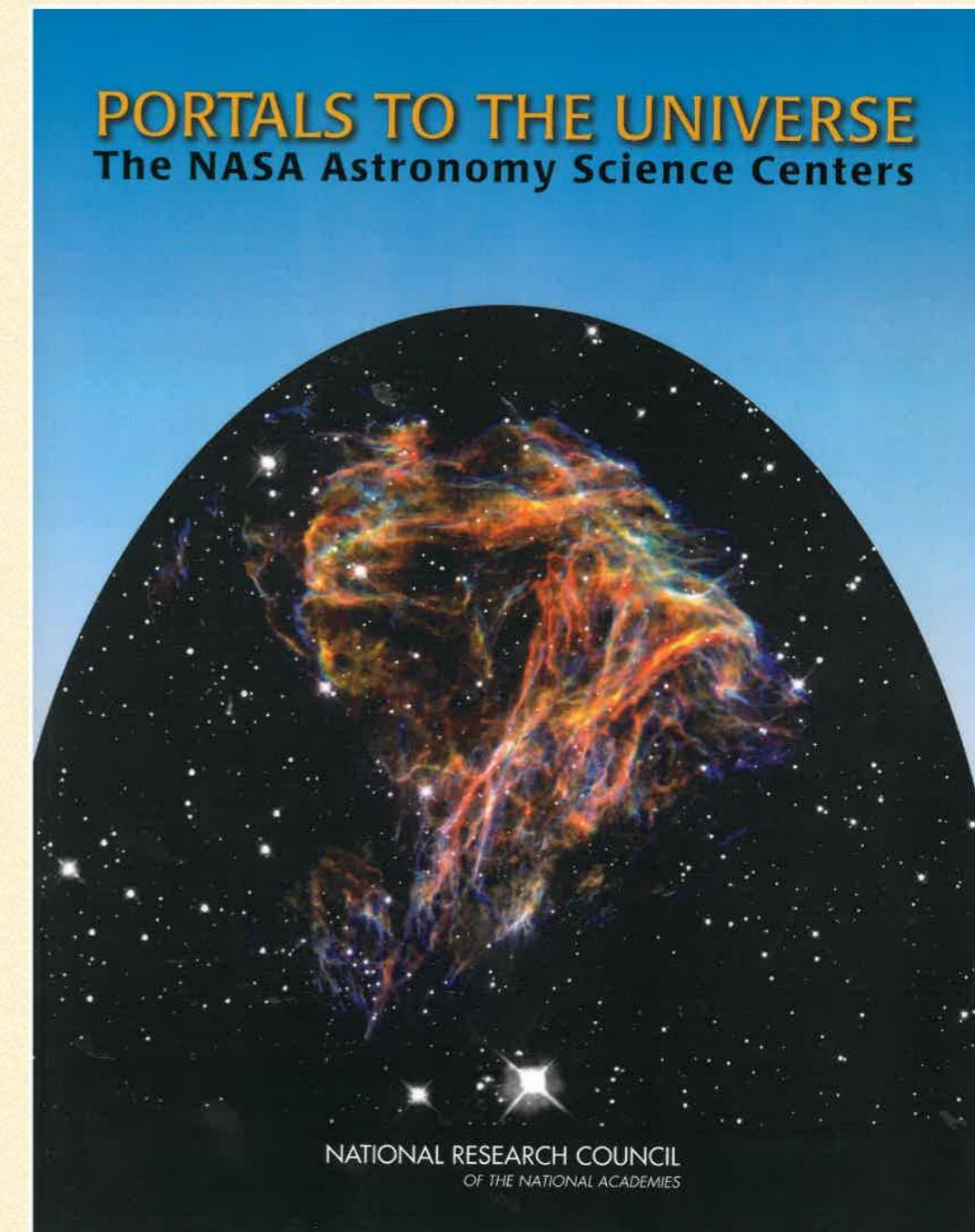
George Helou, Caltech

NRC OIR System Committee  
Irvine, 12 October, 2014

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# PORTALS TO THE UNIVERSE

- Portals to the Universe: The NASA Astronomy Science Centers, NRC Report (2007):
- Science Centers can best process, store and disseminate their data if they maintain mission expertise at the archive centers for the long-term support of active users
- Successful research using archival data sets is dependent on the resident expertise and corporate memory that reside at the science centers.



# QUESTIONS & ANSWERS- I

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- **What is the appropriate ratio of software/hardware cost? What's the key expense?**
- These numbers vary with project and phase
- The key expense is science expertise & user support ( $>1/3$  of total), then software ( $<1/2$ ) and hardware (5-15%)

# QUESTIONS & ANSWERS-2

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- Discuss raw data vs. pipeline data, and the need for pipelines to be done by experts on the particular instruments.
- Without science-ready data, an archive will eventually become unusable and therefore useless. Therefore pipelines are essential
- Extensive scientific familiarity with the instrument is critical in guiding the development of science-ready data
- Expertise needs to be captured in software, reduced data (multiple products!), and documentation to ensure long-term usefulness

# QUESTIONS & ANSWERS-3

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- There's a lack of standardization for ground-based data because of the need to reprocess, specific to different instruments, different conditions, different science goals. How will this intermediate-level processing be possible? Where will it occur? (that is, data will be too large to download on individual computers for reprocessing). Is a coordinated effort needed – data centers, etc.?
- The problem becomes tractable by limiting archive-supported instruments, modes, conditions. Surveys are the obvious extreme case to preserve, but support is not viable at the extreme case of idiosyncratic observations
- The experience gained from archiving space data can contribute critical insights and lessons to ground-based archiving, e.g. instrument operations.  
Effective data archiving starts with instrument design

# QUESTIONS & ANSWERS-4

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- **How/where will archiving take place?**
- Policy addresses the goals and requirements; the answer to how/where depends more on practicalities than on policy.
- The requirements are for data integrity, usability, longevity, etc. Implementing those may be best done by a distributed team, as for the Keck Observatory Archive, leveraging diverse expertise

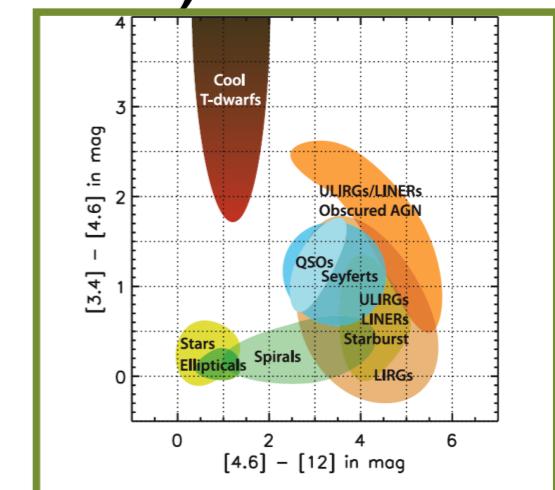
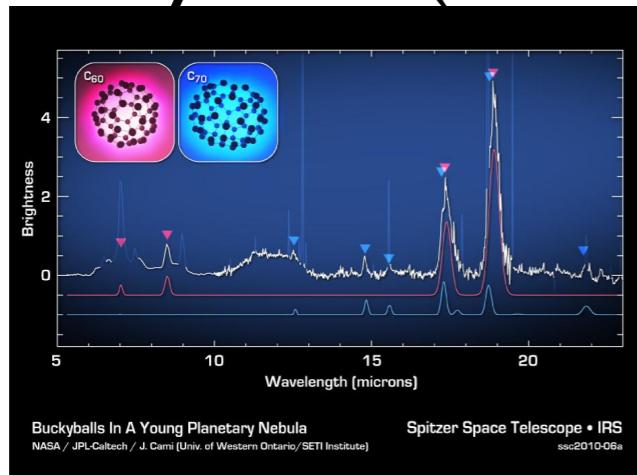
# BACKGROUND: IPAC ARCHIVES

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- IPAC is a multi-mission science and archive center, interfacing projects to community to enable science beyond the expected
- The NASA center of science, data and operations expertise for IR-submm astrophysics, and of Exoplanet Science, especially **IRSA** (InfraRed Science Archive), **NED** (NASA Extragalactic Database), **Exoplanet Archive**, and **Keck Observatory Archive**
- Also supporting **LCOGT** (archive), **PTF/ZTF** (processing and archive), **LSST** (Science User Tools)

# InfraRed Science Archive at IPAC

- Provides efficient access to NASA mission data
  - Began with IRAS and 2MASS, leading into the decade of IR missions
- IRSA serves “Heritage Archives” from Spitzer and other observatories, plus a unique array of all-sky IR surveys covering 20 bands from 1  $\mu$ m to 10 mm
- Supports many research communities from Solar System (NEOWISE) to Cosmology (Planck)



**EXTRAGALACTIC DATABASE NED NASA/IPAC**

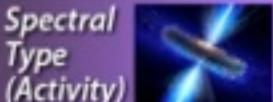
**Overview - What is NED?**  
*THE hub for multi-wavelength research on extragalactic objects*

**Unified multi- $\lambda$  DB and Knowledgebase**

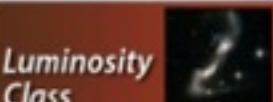
$f(\lambda)$ :

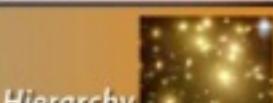
- Name
- $(\alpha, \delta)$
- Redshift
- $D_{\text{Mpc}}$
- Flux
- Diameter
- Attributes
- References
- Notes
- Images
- Spectra
- $A_{\lambda}$

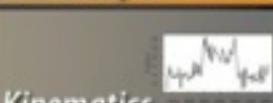
**Galaxy Morphology** 

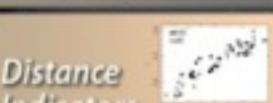
**Spectral Type (Activity)** 

**Radio Morphology** 

**Luminosity Class** 

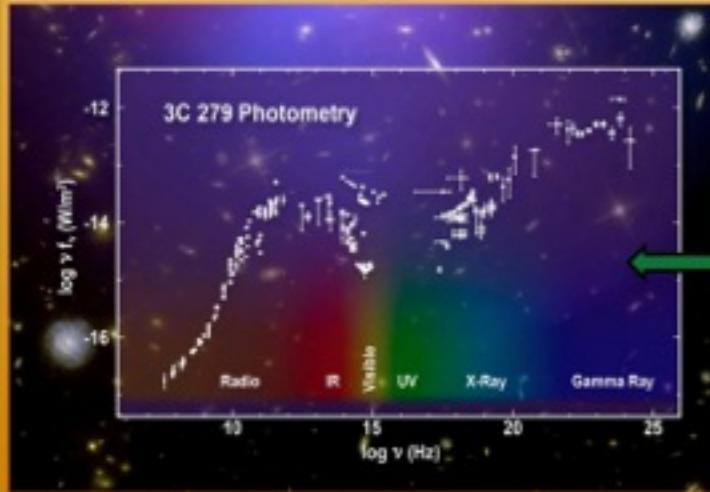
**Hierarchy** 

**Kinematics** 

**Distance Indicators** 

• Standard Formats & Units  
 • Velocity Corrections  
 • Cosmological Corrections  
 • Distances, Luminosities, etc.

• Best available coordinates via any alias: Name Resolver  
 • Spectral Energy Distributions



**Astrophysics Missions**

**E-Literature**

**Large Sky Surveys**

- Research scientists and educators
- Observation and mission planning



- Access via Web, email & computer
- VO program interfaces





## NEW! PI Access to LWS Data and Public Access to ESI Data Now Available: [Data Access](#)

Funded by [NASA](#), the *Keck Observatory Archive (KOA)* is a joint development between the [W. M. Keck Observatory](#) (WMKO) located in Waimea, Hawaii and the [NASA Exoplanet Science Institute](#) (NExScl) located in Pasadena, California. Currently, KOA archives data taken with [HIRES](#), [NIRSPEC](#), [NIRC2](#), [LRIS](#), [KI \(Keck Interferometer\)](#), [MOSFIRE](#), [DEIMOS](#), [ESI](#), [OSIRIS](#), [NIRC](#), and [LWS](#).

The process by which data are archived is as follows:

- Locate all FITS files taken in the previous 24 hour period
- Determine the validity and data integrity of each file
- Add [meta-keywords](#) to the FITS header
- Archive the metadata ([native](#) + [meta-keywords](#)) for each FITS header
- Archive the data for each FITS file
- Archive ancillary data associated with each night's data  
Ancillary data include Keck nightly plots of weather and image quality trends and, when available, all-sky images and photometric data

The physical location of the archive is at [NExScl](#). Both metadata and raw FITS files are transferred via the internet from WMKO to NExScl, and are available through a [User Interface](#) within hours of the end of observations. Both the data files and the metadata are subject to a [proprietary period policy](#). During the proprietary period, program PIs will be able to access their data via a [user login/password protected site](#). The data are released to the public upon expiration of the proprietary period.